2022 School on Electron-Phonon Physics from First Principles 13-19 June 2022, Austin, TX

Hands-On Tutorial Self-trapped polarons with EPW

Introduction

In this session, you will learn to perform calculations of self-trapped polarons with the EPW code. In this short introduction, we briefly present the main concepts and equations relevant for this tutorial. The complete theory and computational method can be found in Phys. Rev. Lett. **122**, 246403 (2019) and Phys. Rev. B **99**, 235139 (2019).

The ground state wave function $\psi(\mathbf{r})$ and atomic displacements $\Delta \tau_{\kappa\alpha p}$ forming a polaron can be found by minimizing the total DFT energy functional of an excess electron added to a crystal, which translates into the solution of the following coupled system of equations:

$$\hat{H}_{\rm KS}^{0}\psi(\mathbf{r}) + \sum_{\kappa\alpha p} \frac{\partial V_{\rm KS}^{0}}{\partial \tau_{\kappa\alpha p}} \Delta \tau_{\kappa\alpha p} \psi(\mathbf{r}) = \varepsilon \,\psi(\mathbf{r}) \,\,, \tag{1}$$

$$\Delta \tau_{\kappa\alpha p} = -\sum_{\kappa'\alpha' p'} (C^0)^{-1}_{\kappa\alpha p,\kappa'\alpha' p'} \int d\mathbf{r} \, \frac{\partial V_{\rm KS}^0}{\partial \tau_{\kappa'\alpha' p'}} \, |\psi(\mathbf{r})|^2 \,. \tag{2}$$

In these expressions, $\tau_{\kappa\alpha p}$ represents the Cartesian coordinate of the atom κ in the unit cell p along the direction α , $C^0_{\kappa\alpha p,\kappa'\alpha'p'}$ is the matrix of interatomic force constants, and $\hat{H}^0_{\rm KS}$ and $V^0_{\rm KS}$ represent the Kohn-Sham Hamiltonian and the self-consistent potential, respectively. The superscript 0 indicates that the quantities are evaluated in the ground state without extra electron. The real space integral is performed over a Born-Von Karman supercell of the crystal containing N_p unit cells. We will refer to ε as the polaron eigenvalue.

By expanding the polaron wave function in terms of the single-particle Kohn-Sham states $\psi_{n\mathbf{k}}$ with eigenvalues $\varepsilon_{n\mathbf{k}}$:

$$\psi(\mathbf{r}) = \frac{1}{\sqrt{N_p}} \sum_{n\mathbf{k}} A_{n\mathbf{k}} \psi_{n\mathbf{k}} , \qquad (3)$$

and the atomic displacements in terms of the phonon eigenmodes $e_{\kappa\alpha,\nu}(\mathbf{q})$ with frequencies $\omega_{\mathbf{q}\nu}$:

$$\Delta \tau_{\kappa\alpha p} = -\frac{2}{N_p} \sum_{\mathbf{q}\nu} B^*_{\mathbf{q}\nu} \left(\frac{\hbar}{2M_\kappa \omega_{\mathbf{q}\nu}}\right)^{1/2} e_{\kappa\alpha,\nu}(\mathbf{q}) e^{i\mathbf{q}\cdot\mathbf{R}_p},\tag{4}$$

where M_{κ} is the mass of the atom κ and \mathbf{R}_p is the lattice vector of the unit cell p, we can transform Eqs. (1) and (2) into a coupled set of equations for the expansion coefficients in reciprocal space:

$$\frac{2}{N_p} \sum_{\mathbf{q}m\nu} B_{\mathbf{q}\nu} g_{mn\nu}^*(\mathbf{k}, \mathbf{q}) A_{m\mathbf{k}+\mathbf{q}} = (\varepsilon_{n\mathbf{k}} - \varepsilon) A_{n\mathbf{k}} , \qquad (5)$$

$$B_{\mathbf{q}\nu} = \frac{1}{N_p} \sum_{mn\mathbf{k}} A^*_{m\mathbf{k}+\mathbf{q}} \frac{g_{mn\nu}(\mathbf{k},\mathbf{q})}{\hbar\omega_{\mathbf{q}\nu}} A_{n\mathbf{k}} .$$
(6)

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The polaron formation energy ΔE_f , defined as the energy required to trap a conduction band state with eigenvalue ε_{CBM} into a localized polaron, can be obtained from the expansion coefficients that solve Eqs. (5) and (6) by:

$$\Delta E_f = \frac{1}{N_p} \sum_{n\mathbf{k}} |A_{n\mathbf{k}}|^2 (\varepsilon_{n\mathbf{k}} - \varepsilon_{\text{CBM}}) - \frac{1}{N_p} \sum_{\mathbf{q}\nu} |B_{\mathbf{q}\nu}|^2 \hbar \omega_{\mathbf{q}\nu} .$$
(7)

We will refer to the first and second terms on the right hand side as the electron and phonon parts of the formation energy, respectively.

From Eqs. (5) and (6) we observe that the necessary ingredients to solve the polaron equations are the Kohn-Sham eigenvalues, phonon frequencies, and electron-phonon matrix elements on the k- and q-grids corresponding to the equivalent Born-Von Karman supercell in which Eqs. (1) and (2) are defined. In order to obtain the formation energy of an isolated polaron, we will need to solve Eqs. (5) and (6) in increasingly denser grids and extrapolate the results to the infinite supercell limit.

Technical note

For this tutorial, we will use a modified version of EPW which includes the polaron module. You can find the source code in the following link:

https://gitlab.com/chaolian/q-e-epw-2022-polaron/

If you want to run this tutorial on your own computer, you can download the compressed tar.bz2 file directly through the following link:

https://gitlab.com/chaolian/q-e-epw-2022-polaron/-/archive/master/q-e-epw-2022-polaron-master.tar.bz2

Extract the files:

(local)\$ tar -xvf q-e-epw-2022-polaron-master.tar.bz2

And compile as usual (https://docs.epw-code.org/doc/DownloadAndInstall.html). In the simplest configuration:

```
(local)$ cd q-e-epw-2022-polaron-master/
(local)$ ./configure
(local)$ make epw
```

We also provide a pre-compiled version of the code in Frontera at:

```
/work2/06868/giustino/EP-SCHOOL/q-e-epw-2022-polaron-master/
```

We recommend to use the pre-compiled version.

You are advised to run the calculations in your scratch folder (\$SCRATCH). To do so, you can go into your scratch folder and copy the tar file which contains all necessary files for this tutorial, and then extract it:

```
$ cd $SCRATCH
$ cp /work2/06868/giustino/EP-SCHOOL/Thu.5.LafuenteLian.tar .
$ tar -xvf Thu.5.LafuenteLian.tar; cd Thu.5.LafuenteLian
```

A quick inspection of the folder will show the two directories containing the files needed in the two exercises, a directory with the pre-completed output files (for reference), and a copy of the pdf of this hands-on:

```
$ ls
exercise1 exercise2 reference Thu.5.LafuenteLian.pdf
```

Note: in this tutorial it will be shown how to obtain all the input files and related files. The folder that you downloaded should be used as a reference or to speed up the process.

Exercise 1

In this exercise you will calculate and analyze the formation of **hole polarons in LiF**. To start, go to the directory of the first exercise:

\$ cd exercise1

The initial steps of the calculation are similar to the ones followed in the previous tutorials. First, we will need to perform an electronic ground state calculation using Quantum Espresso, and a phonon calculation using the PHonon code. The corresponding input files (lif.scf.in and lif.ph.in) are the following:

```
___
&control
                    = 'scf'
   calculation
   prefix
                    = 'lif'
                    = './'
   pseudo_dir
                    = './'
    outdir
    tprnfor
                    = .true.
    tstress
                    = .true.
/
&system
                    = 2
    ibrav
    celldm(1)
                    = 7.67034
                    = 2
   nat
                    = 2
   ntyp
    ecutwfc
                    = 80.0
/
&electrons
                    = 1.0d - 12
    conv_thr
ATOMIC_SPECIES
Li 6.941 Li.upf
F 18.9984 F.upf
```

ATOMIC_POSITIONS crystal Li 0.0000 0.0000 0.0000 F 0.5000 0.5000 0.5000 K_POINTS automatic 6 6 6 0 0 0

&inputph	
prefix	= 'lif'
outdir	= './'
epsil	= .true.
zeu	= .true.
ldisp	= .true.
fildyn	= 'lif.dyn
fildvscf	= 'dvscf'
tr2_ph	= 1.0d-14
nq1	= 6,
nq2	= 6,
nq3	= 6
/	

You can use the following submission job (submit1.sh) to perform the calculation in Frontera:

```
#!/bin/bash
# Job name
                          # Queue (partition) name
                          # Total # of nodes
#SBATCH --ntasks-per-node 24
#SBATCH -t 01:00:00
                          # Run time (hh:mm:ss)
#SBATCH -A EPSchool2022
#SBATCH --reservation=EPSchoolDay4
# echo loaded modules, current directory, and starting time
module list
pwd
date
# export the path which contains executable file
export PATHQE=/work2/06868/giustino/EP-SCHOOL/q-e-epw-2022-polaron-master/
# Launch MPI code...
# Total # of parallel tasks
MPIOPT="-np "$SLURM_NTASKS
# kpoint parallel groups
KPTPRL="-npool 8"
```

```
# run jobs
# scf calculation
ibrun ${MPIOPT} $PATHQE/bin/pw.x ${KPTPRL} -input lif.scf.in > lif.scf.out
# ph calculation
ibrun ${MPIOPT} $PATHQE/bin/ph.x ${KPTPRL} -input lif.ph.in > lif.ph.out
# echo finishing time
date
```

▶ Run a self-consistent calculation and a phonon calculation on a homogeneous $6 \times 6 \times 6$ k and q-point grid.

\$ sbatch submit1.sh

The calculation should take about 5 minutes to complete.

The keyword fildvscf tells the code to write on file the change of the self-consistent potential due to phonon perturbations, $\partial_{\mathbf{q}\nu} V^{\mathrm{scf}}$, that is needed to compute the electron-phonon matrix elements. In the output file lif.dyn0, you can find the list of 16 irreducible \mathbf{q} points in the Brillouin Zone (IBZ). If you type ls, you can see the lif.dynX files containing the dynamical matrix for each irreducible \mathbf{q} point. The dvscf files are all named lif.dvscf1 and are located inside the _ph0/lif.q_X/ folders, except for the one corresponding to the first \mathbf{q} point (Γ) that is located in _ph0/.

► Gather the .dyn and .dvscf files into a new save/ directory which EPW will read.

The files in _ph0/lif.phsave/ containing the displacement patterns are also needed. This can easily be done using the pp.py python script which is already included in the EPW distribution:

\$ python3 /work2/06868/giustino/EP-SCHOOL/q-e-epw-2022-polaron-master/EPW/bin/pp.py

The script will ask you to prompt the prefix of your calculation (in this case lif):

```
Enter the prefix used for PH calculations (e.g. diam) lif
```

 \blacktriangleright Run a non self-consistent calculation on a full homogeneous **k**-point grid, and an EPW calculation to obtain the electron-phonon matrix elements in the Wannier representation.

To proceed with the calculation, we need to obtain the Wannier functions corresponding to the valence band manifold using wannier90 as an internal library, and then calculate the electron-phonon matrix elements in the Wannier representation with EPW.

First we need to perform a non self-consistent calculation on a full homogeneous $6 \times 6 \times 6$ k grid (lif.nscf.in):

```
--
&control
calculation = 'bands'
prefix = 'lif'
```

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```
pseudo_dir
                = './'
   outdir = './'
/
&system
   ibrav
                = 2
   celldm(1) = 7.67034
                 = 2
   nat
   ntyp
                 = 2
   ecutwfc
                = 80.0
   nbnd
                 = 15
/
&electrons
   conv_thr = 1.0d-12
/
ATOMIC_SPECIES
Li 6.941 Li.upf
F 18.9984 F.upf
ATOMIC_POSITIONS crystal
Li 0.0000 0.0000 0.0000
F 0.5000 0.5000 0.5000
K_POINTS crystal
216
0.0000000 0.0000000 0.0000000 4.629630e-03
0.00000000 0.00000000 0.166666667 4.629630e-03
0.0000000 0.0000000 0.3333333 4.629630e-03
 . . .
```

Note: The **k**-point list under K_POINTS crystal in the file above is for illustrative purposes and is not complete. The complete positive-definite homogeneous $6 \times 6 \times 6$ k-point between 0 and 1 can be generated by using the script of kmesh.pl included in the wannier90 package:

\$ /work2/06868/giustino/EP-SCHOOL/q-e-epw-2022-polaron-master/wannier90-3.1.0/utility/kmesh.pl 6 6 6

You can find the complete list in the input file lif.nscf.in. Next, we need to prepare the EPW input file to calculate the electron-phonon matrix elements in the Wannier representation (lif.epw1.in):

_____ &inputepw prefix = 'lif' outdir = './' elph = .true. epwwrite = .true. epwread = .false. lpolar = .true. nbndsub = 3

```
dvscf_dir
                = './save/'
 bands_skipped = 'exclude_bands = 1:2, 6:15'
  wannierize
                = .true.
 num_iter
                = 500
 iprint
                = 2
               = 'F:p'
 proj(1)
 wannier_plot = .true.
 wannier_plot_supercell = 6 6 6
                = 6
 nk1
 nk2
                = 6
 nk3
                = 6
                = 6
 nq1
 nq2
                = 6
                = 6
 nq3
 band_plot
                = .true.
 filkf
                = './path.kpt'
                = './path.kpt'
 filqf
/
```

As you can observe in the input file, we will consider three Wannier functions corresponding to the three bands forming the isolated valence band manifold below the Fermi level. In accordance with their orbital character, we will use three p-orbitals centered at the F-atom as the initial projections. You can now submit the calculation:

\$ sbatch submit2.sh

Note: The submission script submit2.sh can be obtained from the previous submit1.sh by replacing the two lines containing ibrun by the following ones:

```
# nscf calculation
ibrun ${MPIOPT} $PATHQE/bin/pw.x ${KPTPRL} -input lif.nscf.in > lif.nscf.out
# epw calculation
ibrun ${MPIOPT} $PATHQE/bin/epw.x -npool 24 -input lif.epw1.in > lif.epw1.out
```

Note: Similar to what you did in the tutorial Tue.5.Lee, you can assess the quality of the Wannierization by comparing the interpolated electron and phonon band structures with the ones calculated using pw.x and matdyn.x, respectively, as well as by checking the decay of the matrix elements in the Wannier representation as contained in the decay.* files.

 \blacktriangleright Interpolate the electron-phonon matrix elements to a fine **k**- and **q**-point grid and solve the self-consistent polaron equations.

First we need to prepare the EPW input file to perform the interpolation and solve the polaron equations (lif.epw2.in):

&i	inputepw					
	prefix	=	']	lif	<u></u> '	
	outdir	=	۰.	./'	I	
	elph	=	. t	ru	1e.	
	epwread	=	. t	τı	ıe.	
	lpolar	=	. t	ru	ıe.	
	nbndsub	=	3	3		
	dvscf_dir	=	۰.	./s	save/'	
	wannierize	=	. f	fal	Lse.	
	plrn			=	.true.	
	restart_plrn			=	.false.	
	type_plrn			=	1	
	init_plrn			=	1	
	init_sigma_plu	rn		=	1.0	
	niter_plrn			=	500	
	conv_thr_plrn			=	1E-4	
	ethrdg_plrn			=	1E-5	
	nk1	=	6			
	nk2	=	6			
	nk3	=	6			
	nq1	=	6			
	nq2	=	6			
	nq3	=	6			
	nkf1	=	4			
	nkf2	=	4			
	nkf3	=	4			
	nqf1	=	4			
	nqf2	=	4			
	nqf3	=	4			
/						

You can observe that we will use a $4 \times 4 \times 4$ k- and q- point grid in this example. The plrn = .true. tag activates the polaron calculations, and the type_plrn = 1 tag indicates that we are dealing with a hole polaron. The self-consistent polaron equations will be initialized with a Gaussian of width $\sigma = 10.0$ bohr (init_sigma_plrn), and the solution will be considered converged when the greatest difference in the magnitude of the atomic displacements between two subsequent iterations is lower than 10^{-3} bohr (conv_thr_plrn)

You can now submit the calculation:

\$ sbatch submit3.sh

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Note: The submission script submit3.sh can be obtained from the previous submit2.sh by replacing the two lines containing ibrun by the following one:

ibrun \${MPIOPT} \$PATHQE/bin/epw.x -npool 24 -input lif.epw2.in > lif.epw2.out

You can analyze the iterative self-consistent process in the output:

Starting	g the self-cons	istent process			
iter	Phonon/eV	Electron/eV	Formation/eV	Error/eV	Eigval/eV
1	-0.2006E+00	-0.3612E+00	0.1606E+00	0.2099E+00	0.5448E+00
2	-0.1313E+01	-0.7182E+00	-0.5943E+00	0.4092E+00	0.2616E+01
3	-0.2141E+01	-0.8041E+00	-0.1337E+01	0.2069E+00	0.3581E+01
4	-0.2244E+01	-0.8043E+00	-0.1439E+01	0.3546E-01	0.3678E+01
5	-0.2244E+01	-0.8043E+00	-0.1440E+01	0.5459E-03	0.3678E+01
6	-0.2244E+01	-0.8044E+00	-0.1440E+01	0.1472E-03	0.3679E+01

```
End of self-consistent cycle
```

All the information related to the energetics of the converged polaron solution is printed at the end of the self-consistent process:

Eigenvalue	(eV):	3.6785057
Phonon part	(eV):	-2.2441201
Electron part	(eV):	0.8043535
Formation Energy	(eV):	-1.4397666

The breakdown of the formation energy in the different terms is briefly explained in the introduction. More details and further discussion can be found in Phys. Rev. B **99**, 235139 (2019).

A quick inspection of the output files will show that several files have been written with the .plrn extension. The most relevant output files for this exercise are Amp.plrn and dtau.plrn, which contain the polaron wave function coefficients in the Wannier basis and the atomic displacements, respectively. These files are needed to postprocess and analyze the polaron solution in more detail.

▶ Plot and analyze polaron wave function.

The post-processing of the polaron solution can be activated by the following input file (lif.epw3.in):

&inputepw			
prefix	= 'lif'		
outdir	= './'		
elph	= .true.		
epwread	= .true.		
lpolar	= .true.		

```
= 3
nbndsub
dvscf_dir = './save/'
wannierize = .false.
plrn
                = .true.
restart_plrn
               = .true.
type_plrn
                = 1
cal_psir_plrn = .true.
interp_Ank_plrn = .true.
interp_Bqu_plrn = .true.
filkf
                 = './path.kpt'
filqf
                = './path.kpt'
             = 6
nk1
nk2
             = 6
nk3
             = 6
nq1
             = 6
nq2
             = 6
             = 6
nq3
nkf1
             = 4
nkf2
             = 4
nkf3
             = 4
             = 4
nqf1
nqf2
             = 4
nqf3
             = 4
```

The input tag restart_plrn = .true. skips the self-consistent solution of the polaron equations and triggers the post-processing of the previously converged solution by reading the Amp.plrn and dtau.plrn files.

The input tag cal_psir_plrn = .true. calls the subroutine to plot the polaron wave function in real space, which is contained in the psir_plrn.xsf file. This file can be plotted with VESTA. First submit the calculation:

\$ sbatch submit4.sh

and then copy the file containing the wave function to your own computer:

\$ scp username@frontera.tacc.utexas.edu:path_to_exercise/psir_plrn.xsf .

As when logging in, enter your TACC password at the password prompt. At the TACC Token prompt, enter your 6-digit code. Then you can visualize the polaron by running VESTA locally on your computer:

(local)\$ VESTA psir_plrn.xsf

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You should obtain something similar to this:



The polaron wave function is plotted as an isosurface, and the red arrows indicate the magnitude and direction of the atomic displacements in the polaronic configuration. You can observe that the $4 \times 4 \times 4$ k-grid transforms to a $4 \times 4 \times 4$ supercell in real space, with periodic boundary conditions. From this figure it is also appreciated that holes form small polarons in LiF, localized in one unit cell, and have the shape of a *p*-orbital centered on a F-atom.

Next we will analyze the polaron wave function coefficients in the single-particle basis of the DFT-Kohn Sham states. This calculation is activated by the interp_Ank_plrn = .true. tag. Similarly, the calculation of the expansion coefficients of the atomic displacements in the phonon basis is activated by interp_Bqu_plrn = .true. The files containing the **k**- and **q**-points in which the polaron coefficients will be interpolated have to be indicated in the filkf and filqf input tags, respectively. In this example we will use a path along the high-symmetry W-L- Γ -X-W-K points, as contained in the path.kpt file.

The output files containing the interpolated coefficients are Ank.band.plrn and Bmat.band.plrn. You can visualize your results by using, for example, the following script (plot_Ank.py):

```
## Script to plot Ank coefficients
import matplotlib as mpl
mpl.use('pdf')
import matplotlib.pyplot as plt
import numpy as np
# change font on mathematical expressions on plots
mpl.rcParams['mathtext.fontset'] = 'cm'
# Read kpath file
kpath = np.loadtxt('path.kpt', skiprows=1)
# Read Ank file
ik, ibnd, ek0, ReAnk, ImAnk, AbsAnk = np.loadtxt('Ank.band.plrn',
```

```
unpack=True, skiprows=1)
# Separate data in different bands
nbnd=np.int(max(ibnd))
maxik=np.int(max(ik))
Ak=np.zeros((maxik,nbnd))
ek=np.zeros((maxik,nbnd))
iklist=np.zeros((maxik,nbnd))
for i in range(maxik):
    for ibnd in range(nbnd):
        Ak[i][ibnd]=AbsAnk[i*nbnd+ibnd]
        ek[i][ibnd]=ek0[i*nbnd+ibnd]
        iklist[i][ibnd]=i
# Get vbm and bandwidth
vbm=max(ek[:,nbnd-1])
bandwidth=vbm-min(ek[:,0])
## Plot bands
f, ax = plt.subplots(figsize=(10,6))
ax.plot(iklist, ek, color='blue')
# Plot Ank
ax.scatter(iklist, ek, 100*Ak, color='gold', edgecolors='gray', alpha=0.8,
           label=r'$|A_{n\mathbb{k}}|$')
# Set tick params etc.
ax.set_ylim(-bandwidth-0.1*bandwidth,0.1*bandwidth)
ax.set_xlim((min(iklist[:,0]),max(iklist[:,0])))
ax.set_xticklabels([])
ax.tick_params(axis='x', color='black', labelsize='0', pad=0, length=0, width=0)
ax.tick_params(axis='y', color='black', labelsize='18', pad=5, length=5, width=1)
ax.set_ylabel(r'$E-E_{\mathrm{VBM}} ~ (\mathrm{eV})$', fontsize=25, labelpad=10)
ax.legend(loc='upper right', fontsize=25)
plt.savefig('Ank.pdf')
#plt.show()
```

Note: You can find a similar script to plot the $B_{q\nu}$ coefficients in the example directory (plot_Bqv.py).

Executing these scripts:

\$ python3 plot_Ank.py
\$ python3 plot_Bqv.py

should generate two .pdf output files (Ank.pdf and Bqv.pdf) containing the corresponding figures. To visualize the plots, copy the output files to your own computer (enter your TACC password and token as

before):

\$ scp username@frontera.tacc.utexas.edu:path_to_exercise/*.pdf .

and open them with your preferred .pdf file reader. For example, in Linux you can type:

(local)\$ evince Ank.pdf
(local)\$ evince Bqv.pdf

Your figures should be similar to:



As you can observe in these results, due to the strongly localized nature of the hole polaron in LiF, its wave function coefficients in the Bloch basis are spread across the entire Brillouin zone (see Eqs. (3) and (4) in the introduction).

► Study the change of the polaron formation energy as a function of the supercell size, and extrapolate results to isolated polaron in an infinite supercell.

In the previous calculation we have seen that the $4 \times 4 \times 4$ k- and q-point grid in momentum space corresponds to a $4 \times 4 \times 4$ supercell in real space. Due to the interaction between replicated polarons in the neighbouring supercells, the formation energy of the polaron will depend on the momentum grid (supercell) size.

You can analyze the dependence of the polaron formation energy on the supercell size by using the following script (submit5.sh):

```
#!/bin/bash
#SBATCH -J myjob
                            # Job name
                             # Queue (partition) name
#SBATCH -p small
                             # Total # of nodes
#SBATCH -N 1
#SBATCH --ntasks-per-node 48
#SBATCH -t 00:30:00
                            # Run time (hh:mm:ss)
#SBATCH -A EPSchool2022
#SBATCH --reservation=EPSchoolDay4
# echo loaded modules, current directory, and starting time
module list
pwd
date
# export the path which contains executable file
export PATHQE=/work2/06868/giustino/EP-SCHOOL/q-e-epw-2022-polaron-master
# Launch MPI code...
echo "#
          nk Eform" > "E_vs_nk.dat"
for i in 6 8 10 12
do
sed -e "30s/.*/ nkf1
                          = $i/" \
   -e "31s/.*/ nkf2
                            = $i/" \
                             = $i/" \
   -e "32s/.*/ nkf3
   -e "33s/.*/ nqf1
                             = $i/" \
   -e "34s/.*/ nqf2
                            = $i/" \
   -e "35s/.*/ nqf3
                             = $i/" \
    "lif.epw4.in" > "lif.epw4.$i.in"
# run jobs
j=$(( 4*$i ))
ibrun -np $j $PATHQE/bin/epw.x -npool $j -input lif.epw4.$i.in > lif.epw4.$i.out
grep 'Formation Energy (eV):' "lif.epw4.$i.out" >> "E_vs_nk.dat"
sed -i "s/.Formation Energy (eV):/ $i/" "E_vs_nk.dat"
done
# echo finishing time
```

date

You can submit this job by:

\$ sbatch submit5.sh

For each of the desired **k**-point grids (in this case 6^3 , 8^3 , 10^3 and 12^3), this script creates a copy of the input file lif.epw4.in with modified nkf and nqf input variables into lif.epw4.*.in, launches the calculation, and copies the resulting formation energy to the E_vs_nk.dat file:

```
# nk Eform
6 -1.3929285
8 -1.6041677
10 -1.7337758
12 -1.8089613
```

You can plot the formation energy as a function of the inverse supercell size (here defined as L^{-1} where L^3 is the supercell volume) and extrapolate to the infite supercell by using, for example, the following script (plot_E_s_nk.py):

```
import matplotlib as mpl
mpl.use('pdf')
import matplotlib.pyplot as plt
from matplotlib.ticker import MultipleLocator,FormatStrFormatter
import numpy as np
# Read E vs nk
Nk, Eform = np.genfromtxt('E_vs_nk.dat', unpack=True)
# Set unit cell volume to convert Nk to inverse supercell size
ucell_volume = 112.8044
# Start figure
fig, ax = plt.subplots(figsize=(12,8))
# Plot Eform
ax.scatter(1/(Nk*ucell_volume**(1/3)), Eform, s=50, marker='o',
           color='darkred', edgecolors='black')
# Perform linear fit and plot
mf, bf = np.polyfit(1/(Nk*ucell_volume**(1/3)), Eform, 1)
xlist = np.linspace(0.0, np.max(1/(Nk*ucell_volume**(1/3))), 100)
print("Extrapolation to isolated polaron formation energy = ", "%.3f" % bf, "eV")
ax.plot(xlist, mf*xlist+bf, '--', color='gray')
# Set tick params etc.
ax.set_xlabel(r'Inverse supercell size ($\mathrm{bohr}^{-1}$)',fontsize=20)
```

Executing this script :

\$ python3 plot_E_vs_nk.py

copying the output file to your own computer:

\$ scp username@frontera.tacc.utexas.edu:path_to_exercise/E_vs_nk.pdf .

and opening the .pdf file locally:

(local)\$ evince E_vs_nk.pdf

You should obtain something similar to this:



In this example, the extrapolation to infinite supercell gives a formation energy of $\Delta E_f = -2.232 \text{ eV}$ for the isolated hole polaron in LiF. Note that the present results are not fully converged (see Phys. Rev. B **99**, 235139 (2019) for fully converged calculation parameters).

Exercise 2

In this exercise you will repeat a similar procedure as before, but this time to calculate and analyze the formation of **electron polarons in LiF**.

Start by moving to the directory for this exercise, and creating symbolic links to the lif.save and save directories generated in the previous exercise, so that we don't need to run the scf, ph, and nscf calculations again:

```
$ cd ../exercise2/
$ ln -s ../exercise1/lif.save/ .
$ ln -s ../exercise1/save/ .
```

You can follow the calculation steps one by one as in Exercise 1, and analyze the intermediate results. Alternatively, you can submit the following script that will perform all the calculations automatically (submit1.sh):

```
#!/bin/bash
#SBATCH -J myjob
                              # Job name
#SBATCH -p small
                              # Queue (partition) name
                              # Total # of nodes
#SBATCH -N 1
#SBATCH --ntasks-per-node 56
#SBATCH -t 01:00:00
                              # Run time (hh:mm:ss)
#SBATCH -A EPSchool2022
#SBATCH --reservation=EPSchoolDay4
# echo loaded modules, current directory, and starting time
module list
pwd
date
# export the path which contains executable file
export PATHQE=/work2/06868/giustino/EP-SCHOOL/q-e-epw-2022-polaron-master
# run jobs
# epw1: g(Re,Rp)
ibrun -np 54 $PATHQE/bin/epw.x -npool 54 -input lif.epw1.in > lif.epw1.out
# epw2: polaron for different nk
echo "#
                   Eform" > "E_vs_nk.dat"
           nk
for i in 6 8 10 12 14
do
sed -e "31s/.*/ nkf1
                              = $i/" \
```

```
-e "32s/.*/ nkf2
                             = $i/" \
                            = $i/" \
   -e "33s/.*/ nkf3
                            = $i/" \
   -e "34s/.*/ nqf1
   -e "35s/.*/ nqf2
                             = $i/" \
                            = $i/" \
   -e "36s/.*/ nqf3
    "lif.epw2.in" > "lif.epw2.$i.in"
# run jobs
j=$(( 4*$i ))
ibrun -np $j $PATHQE/bin/epw.x -npool $j -input lif.epw2.$i.in > lif.epw2.$i.out
grep 'Formation Energy (eV):' "lif.epw2.$i.out" >> "E_vs_nk.dat"
sed -i "s/.Formation Energy (eV):/ $i/" "E_vs_nk.dat"
done
```

Note: Analyze the input file lif.epw2.in. You will see that in this case we consider one conduction band for the Wannierization, and that the **electron** polaron calculation is set by the type_plrn = -1 input tag.

▶ Study the change of the electron polaron formation energy as a function of the supercell size.

You can submit the calculation with:

\$ sbatch submit1.sh

Note: This calculation should take around 10 minutes to complete in Frontera.

The dependence of the polaron formation energy with respect to the momentum grid (supercell) size in this case is $(E_vs_nk.dat)$:

#	nk	Eform
	6	0.000015
	8	0.000043
	10	0.0000491
	12	-0.0572563
	14	-0.1041795

Exercise: Why do we obtain finite formation energies only for supercell sizes larger than $12 \times 12 \times 12$?

Similarly as before, we can postprocess and analyze the converged polaron solution by studying its wave function in real space, and the distribution of the wave function coefficients in reciprocal space. We will analyze the polaron solution obtained in the last $14 \times 14 \times 14 \times 14$ k-point grid calculation:

\$ sbatch submit2.sh

Note: The calculation of the polaron wave function in the $14 \times 14 \times 14$ real space supercell should take around 8 min to complete in Frontera.

Note: You can obtain the lif.epw3.in input file by modifying lif.epw2.14.in in a similar way as we have done before for lif.epw3.in in Exercise 1. The input tag step_wf_grid_plrn = 2 reduces the amount of real-space points in which the polaron wave function is computed, and it is only used here to reduce the computational time. You can obtain the submit2.sh submission file by copying submit4.sh from ../exercise1/

The distribution of the polaron wave function coefficients along the high-symmetry W-L- Γ -X-W-K points should be similar to:

Note: You can obtain plot the polaron wave function coefficients by copying plot_Ank.py and plot_Bqv.py from .../exercise1/, and modifying the following lines in plot_Ank.py: ax.set_ylim(-0.1*bandwidth,bandwidth+0.1*bandwidth) ax.set_ylabel(r'\$E-E_{\mathrm{CBM}} ~ (\mathrm{eV})\$', fontsize=25, labelpad=10)

Note: The small oscillations that might appear in the plot for the $B_{q\nu}$ coefficients are a spurious effect due to the fact that the $14 \times 14 \times 14$ k-point grid in which the polaron equations are solved is not large enough. They vanish for denser k-point grids.



And the real space plot of the polaron wave function should look similar to:



As you can observe in these results, electrons in LiF form **large** polarons in real space, which correspond to wave function coefficients localized around the conduction band bottom.