

2022 School on Electron-Phonon Physics from First Principles

Calculations of superconducting properties with EPW

Hands-on session (Wed.7)

Hands-on based on QE-v7.1 and EPW-v5.5

Exercise 1

In this example we are going to calculate the superconducting properties of fcc Pb by solving the isotropic Migdal-Eliashberg equations. The theory related to this tutorial can be found in [Phys. Rev. B 87, 024505 \(2013\)](#).

```
$ cd $SCRATCH
$ mkdir EP-SCHOOL ; cd EP-SCHOOL
$ cp /work2/06868/giustino/EP-SCHOOL/Wed.7.Margine.tar .
$ tar -xvf Wed.7.Margine.tar; cd Wed.7.Margine/exercise1
```

► 1st step: Run a self-consistent calculation on a homogeneous $12 \times 12 \times 12$ **k**-point grid and a phonon calculation on a homogeneous $3 \times 3 \times 3$ **q**-point grid using the following jobscript (`job.ph`) and input files (`scf.in` and `ph.in`):

Note: The `ecutwfc` needs to be much larger for real calculations.

```
$ cd phonon
$ sbatch job.ph
```

```
#!/bin/bash                                                                                                     job.ph
#SBATCH -J job.ph          # Job name
#SBATCH -N 1              # Total # of nodes
#SBATCH --ntasks-per-node 8
#SBATCH -t 01:00:00      # Run time (hh:mm:ss)
#SBATCH -A EPSchool2022
#SBATCH -p small
#SBATCH --reservation=EPSchoolDay3

# Launch MPI code...
export PATHQE=/work2/06868/giustino/EP-SCHOOL/q-e

ibrun $PATHQE/bin/pw.x -nk 4 -in scf.in > scf.out
ibrun $PATHQE/bin/ph.x -nk 4 -in ph.in > ph.out
```

```
--                                                                                                       scf.in
&control
  calculation = 'scf'
  restart_mode = 'from_scratch',
  prefix      = 'pb',
  pseudo_dir  = '.././pseudo/',
  outdir      = './'
/
&system
 ibrav      = 2,
celldm(1)  = 9.27,
nat        = 1,
```

```

ntyp          = 1,
ecutwfc       = 30.0
occupations   = 'smearing',
smearing      = 'mp',
degauss       = 0.025
/
&electrons
  diagonalization = 'david'
  mixing_beta     = 0.7
  conv_thr        = 1.0d-12
/
ATOMIC_SPECIES
Pb 207.2 pb_s.UPF
ATOMIC_POSITIONS
Pb 0.00 0.00 0.00
K_POINTS {automatic}
12 12 12 0 0 0

```

```

--
&inputph
prefix = 'pb',
fildyn = 'pb.dyn.xml',
fildvscf = 'dvscf',
tr2_ph = 1.0d-17
ldisp = .true.,
nq1 = 3,
nq2 = 3,
nq3 = 3

```

ph.in

During the run, notice the irreducible (IBZ) q-point grid:

```

Dynamical matrices for ( 3, 3, 3) uniform grid of q-points
( 4 q-points):
  N      xq(1)      xq(2)      xq(3)
  1  0.000000000  0.000000000  0.000000000
  2 -0.333333333  0.333333333 -0.333333333
  3  0.000000000  0.666666667  0.000000000
  4  0.666666667 -0.000000000  0.666666667

```

► 2nd step: Gather the `.dyn`, `.dvscf`, and `patterns` files into a new save directory using the `pp.py` python script.

```
$ python3 /work2/06868/giustino/EP-SCHOOL/q-e/EPW/bin/pp.py
```

The script will ask you to provide the prefix of your calculation (here "pb").

► 3rd step: Do a non self-consistent calculation on a $3 \times 3 \times 3$ **homogeneous** and Γ -**centered grid between [0,1]** in crystal coordinates and an EPW calculation for the superconducting properties using the following inputs and jobscrip:

Note 1: The homogeneous k grid for the non self-consistent calculations can be generated using the script [kmesh.pl](#)

```
$ /work2/06868/giustino/EP-SCHOOL/q-e/wannier90-3.1.0/utility/kmesh.pl 3 3 3
```

Note 2: A non self-consistent calculation requires the charge density found from a previous self-consistent run with `pw.x`. In the jobscrip `job.epw1` you can see that a self-consistent calculation is run first with the same `scf.in` file used in the `phonon` directory. Alternatively, one can make the `pb.save` directory and copy there the files from `phonon/pb.save` For this in `job.epw1` you need to comment the line

```
#ibrun $PATHQE/bin/pw.x -nk 8 -in scf.in > scf.out
```

and uncomment the following three lines

```
mkdir pb.save
cp ../phonon/pb.save/charge-density.dat pb.save/
cp ../phonon/pb.save/data-file-schema.xml pb.save/
```

Note 3: EPW calculations with `ephwrite = .true.` require that the fine `k` or `q` grids are commensurate, i.e., `nkf1`, `nkf2`, `nkf3` to be multiple of `nqf1`, `nqf2`, `nqf3`.

```
$ cd ../epw
$ sbatch job.epw1
```

```

                                                                    job.epw1
#!/bin/bash
#SBATCH -J job.epw1           # Job name
#SBATCH -N 1                 # Total # of nodes
#SBATCH --ntasks-per-node 8
#SBATCH -t 01:00:00         # Run time (hh:mm:ss)
#SBATCH -A EPSchool2022
#SBATCH -p small
#SBATCH --reservation=EPSchoolDay3

# Launch MPI code...
export PATHQE=/work2/06868/giustino/EP-SCHOOL/q-e

ibrun $PATHQE/bin/pw.x -nk 8 -in scf.in > scf.out
#alternatively to re-run a scf calculation copy files from ../phonon/pb.save
#mkdir pb.save
#cp ../phonon/pb.save/charge-density.dat pb.save/
#cp ../phonon/pb.save/data-file-schema.xml pb.save/

ibrun $PATHQE/bin/pw.x -nk 8 -in nscf.in > nscf.out
ibrun $PATHQE/bin/epw.x -nk 8 -in epw1.in > epw1.out

```

```

                                                                    nscf.in
--
&control
  calculation = 'nscf',
  restart_mode = 'from_scratch',
  prefix      = 'pb',
  pseudo_dir  = '../pseudo/',
  outdir      = './',
  verbosity   = 'high'
/
&system
  ibrav      = 2,
  celldm(1)  = 9.27,
  nat        = 1,
  ntyp       = 1,
  ecutwfc    = 30.0,
  occupations = 'smearing',
  smearing   = 'mp',
  degauss    = 0.025,
  nbnd       = 10,
/
&electrons
  diagonalization = 'david'
  mixing_beta     = 0.7
  conv_thr        = 1.0d-12
/
ATOMIC_SPECIES
Pb 207.2 pb_s.UPF
ATOMIC_POSITIONS crystal
Pb 0.000000000 0.000000000 0.000000000
K_POINTS crystal
27
0.00000000 0.00000000 0.00000000 3.703704e-02
0.00000000 0.00000000 0.33333333 3.703704e-02
...

```

```

--                                                                 epw1.in
&inputepw
  prefix      = 'pb',
  outdir      = './'
  dvscf_dir   = './phonon/save' ! directory where .dyn, .dvscf and prefix.phsave/patterns.xx.yy
                                ! files obtained from phonon calculation are stored

  ep_coupling = .true.          ! run e-ph coupling calculation
  elph        = .true.          ! calculate e-ph coefficients
  epwwrite    = .true.          ! write e-ph matrices in Wann representation
  epwread     = .false.         ! read e-ph matrices from 'prefix.epmatwp' file

  wannierize  = .true.          ! calculate Wannier functions using W90 library
  nbndsub     = 4                ! number of Wannier functions to utilize
  bands_skipped = 'exclude_bands = 1-5' ! number of bands skipped during wannierization

  num_iter    = 300
  dis_froz_min = -3
  dis_froz_max = 13.5
  proj(1)     = 'Pb:sp3'
  wdata(1)    = 'bands_plot = .true.'
  wdata(2)    = 'begin kpoint_path'
  wdata(3)    = 'G 0.00 0.00 0.00 X 0.00 0.50 0.50'
  wdata(4)    = 'X 0.00 0.50 0.50 W 0.25 0.50 0.75'
  wdata(5)    = 'W 0.25 0.50 0.75 L 0.50 0.50 0.50'
  wdata(6)    = 'L 0.50 0.50 0.50 K 0.375 0.375 0.75'
  wdata(7)    = 'K 0.375 0.375 0.75 G 0.00 0.00 0.00'
  wdata(8)    = 'G 0.00 0.00 0.00 L 0.50 0.50 0.50'
  wdata(9)    = 'end kpoint_path'
  wdata(10)   = 'bands_plot_format = gnuplot'

  fsthick     = 0.4              ! Fermi window thickness [eV]
  degaussw    = 0.1             ! smearing in energy-conserving delta functions in [eV]
  degausssq   = 0.05           ! smearing for sum over q in the e-ph coupling in [meV]

  ephwrite    = .true.          ! write ephmatXX, egnv, freq, and ikmap files in prefix.ephmat directory
  eliashberg  = .true.          ! calculate Eliashberg spectral function

  liso        = .true.          ! solve isotropic ME eqs.
  limag       = .true.          ! solve ME eqs. on imaginary axis
  lpade       = .true.          ! solve ME eqs. on real axis using Pade approximants
  lacon       = .true.          ! analytic continuation of ME eqs. from imaginary to real axis

  nsiter      = 500             ! number of self-consistent iterations when solving ME eqs.
  npade       = 40              ! percentage of Matsubara points used in Pade continuation.
  conv_thr_iaxis = 1.0d-3       ! convergence threshold for solving ME eqs. on imaginary axis
  conv_thr_racon = 1.0d-3       ! convergence threshold for solving ME eqs. on real axis

  wscut       = 0.1             ! upper limit over Matsubara freq. summation in ME eqs on imag.axis in [eV]
  muc         = 0.1             ! effective Coulomb potential used in the ME eqs.
  temps       = 0.3 0.9 1.5 2.1 2.7 3.3 3.9 4.2 4.4 4.5 4.6 4.65 ! temperatures at which ME eqs.
                                ! are solved [equally spaced temperature points can also be used: see epw2.in]

  nk1 = 3
  nk2 = 3                ! dimensions of coarse electronic grid
  nk3 = 3

  nq1 = 3
  nq2 = 3                ! dimensions of coarse phonon grid
  nq3 = 3

  mp_mesh_k = .true.      ! use irreducible electronic fine mesh
  nkf1 = 18
  nkf2 = 18                ! dimensions of fine electronic grid
  nkf3 = 18

  nqf1 = 18
  nqf2 = 18                ! dimensions of phonon grid
  nqf3 = 18

```

/

With the above input, we are instructing EPW to:

- Fourier-transform the electron-phonon matrix elements from a coarse $3 \times 3 \times 3$ to a dense $18 \times 18 \times 18$ \mathbf{k}/\mathbf{q} -point grids.

```
Using uniform q-mesh:  18  18  18
Size of q point mesh for interpolation:      5832
Using uniform MP k-mesh:  18  18  18
Size of k point mesh for interpolation:      390
Max number of k points per pool:           50
```

- Pre-compute the \mathbf{q} -points that fall within the `fsthick` window. If at a specific \mathbf{q} -point at least one $\mathbf{k} + \mathbf{q}$ eigenvalue falls within the user-defined `fsthick`, then the \mathbf{q} -point is selected.

```
Number selected, total      100      107
Number selected, total      200      215
.....
Number selected, total      5500     5736
We only need to compute    5579 q-points
```

- Write on disk in the `pb.ephmat` directory the: (1) `ephmatXX` files (one per CPU) containing the electron-phonon matrix elements within the Fermi window (`fsthick`) on the dense \mathbf{k} and \mathbf{q} grids, (2) `freq` file containing the phonon frequencies on the dense \mathbf{q} grid, (3) `egnv` file containing the eigenvalues within the Fermi window on the dense \mathbf{k} grid, and (4) `ikmap` file containing the index of the \mathbf{k} -points on the dense (irreducible) grid within the Fermi window. All these files are produced by setting `ephwrite = .true.`. The files are unformatted and required for solving the Migdal-Eliashberg equations.

```
Nr. of irreducible k-points on the uniform grid:      195
```

```
Finish mapping  $\mathbf{k} + \text{sign} \times \mathbf{q}$  onto the fine irreducible k-mesh and writing .ikmap file
```

```
Nr irreducible k-points within the Fermi shell =      30 out of      195
```

```
Progression iq (fine) =      100/      5579
Progression iq (fine) =      200/      5579
.....
Progression iq (fine) =      5500/      5579
      Fermi level (eV) =      0.117577170439830D+02
DOS(states/spin/eV/Unit Cell) =      0.296296405828119D+00
      Electron smearing (eV) =      0.100000000000000D+00
      Fermi window (eV) =      0.400000000000000D+00
```

```
Finish writing .ephmat files
```

- Solve the isotropic Migdal-Eliashberg equations on the imaginary frequency axis. This is achieved by setting the keywords `eliashberg = .true.`, `liso = .true.`, and `limag = .true.` in the input file. The equations are solved self-consistently for each temperature value specified in the input file. The calculation at each temperature ends when either the converge threshold (`conv_thr_iaxis`) or the maximum number of iterations (`nsiter`) is reached.

Note 1: If at a specific temperature the maximum number of iterations is reached without achieving convergence, the code will stop and not move to the next temperature in the list.

Note 2: Because the electron-phonon matrix elements do not depend on the temperature at which the Migdal-Eliashberg equations are solved, they can be reused in subsequent EPW calculations at different temperatures. This is the reason why `ephmatXX` files are saved in the `pb.ephmat` directory.

The isotropic Migdal-Eliashberg equations take the following form:

$$\begin{aligned}
 Z(i\omega_j) &= 1 + \frac{\pi T}{\omega_j} \sum_{j'} \frac{\omega_{j'}}{\sqrt{\omega_{j'}^2 + \Delta(i\omega_j)}} \lambda(\omega_j - \omega_{j'}) \\
 Z(i\omega_j) \Delta(i\omega_j) &= \pi T \sum_{j'} \frac{\Delta(i\omega_{j'})}{\sqrt{\omega_{j'}^2 + \Delta^2(i\omega_{j'})}} [\lambda(\omega_j - \omega_{j'}) - \mu_c^*]
 \end{aligned} \tag{1}$$

The semiempirical Coulomb parameter μ_c^* is provided as an input variable `muc` in the EPW calculation. The isotropic electron-phonon coupling strength $\lambda(\omega_j)$ entering in Eqs. (1) is defined as:

$$\lambda(\omega_j) = \frac{1}{N_F} \sum_{nm\nu} \int_{\Omega_{\text{BZ}}} \frac{d\mathbf{k}}{\Omega_{\text{BZ}}} \int_{\Omega_{\text{BZ}}} \frac{d\mathbf{q}}{\Omega_{\text{BZ}}} |g_{mn\nu}(\mathbf{k}, \mathbf{q})|^2 \frac{2\omega_{\mathbf{q}\nu}}{\omega_j^2 + \omega_{\mathbf{q}\nu}^2} \delta(\epsilon_{n\mathbf{k}} - \epsilon_F) \delta(\epsilon_{m\mathbf{k}+\mathbf{q}} - \epsilon_F) \tag{2}$$

While the calculation is running, notice in the `epw1.out` file the different steps a full EPW run goes through. Once the interpolation on the fine mesh is finished, the code writes and reads the files required for solving the Migdal-Eliashberg equations and then proceeds with solving the equations at the specified temperatures.

```

=====
Solve isotropic Eliashberg equations
=====
Finish reading freq file
.....

1 bands within the Fermi window

Finish reading egnv file

Max nr of q-points =          956

Finish reading ikmap files

Start reading .ephmat files

Finish reading .ephmat files

a2f file is not found to estimate initial gap: calculationg a2f files

Finish reading a2f file

Electron-phonon coupling strength =    1.4873591

Estimated Allen-Dynes Tc =    3.589807 K for muc =    0.10000

Estimated w_log in Allen-Dynes Tc =    2.735667 meV

```

```

Estimated BCS superconducting gap =      0.544448 meV

WARNING WARNING WARNING

The code may crash since tempmax =      4.650 K is larger than Allen-Dynes Tc =      3.590 K

temp( 1) =      0.30000 K

Solve isotropic Eliashberg equations on imaginary-axis

Total number of frequency points nsiw( 1) =      616
Cutoff frequency wscut =      0.1001

iter      ethr      znormi      deltai [meV]
1  3.197751E+00  2.354591E+00  6.794531E-01
2  1.552188E-01  2.336374E+00  7.345734E-01
.....
8  1.751508E-04  2.285988E+00  8.740011E-01
Convergence was reached in nsiter =      8

```

- Perform an analytic continuation of the solutions from the imaginary frequency axis to the real frequency axis. The analytic continuation can be done using Padé approximants (`lpade = .true.`) or an iterative procedure (`lacon = .true.`). The iterative procedure is performed self-consistently until either the converge threshold (`conv_thr_racon`) or the maximum number of iterations (`nsiter`) is reached.

Note: If at a specific temperature the maximum number of iterations is reached without achieving convergence, the code will stop and not move to the next temperature in the list.

```

Padé approximant of isotropic Eliashberg equations from imaginary-axis to real-axis
Cutoff frequency wscut =      0.1000

```

```

pade      Re[znorm]      Re[delta] [meV]
246  2.286410E+00  8.742188E-01
Convergence was reached for N =      246 Padé approximants

```

```

raxis_pade      :      0.02s CPU      0.05s WALL (      1 calls)

```

Analytic continuation of isotropic Eliashberg equations from imaginary-axis to real-axis

```

Total number of frequency points nsw =      5000
Cutoff frequency wscut =      0.1000

```

```

iter      ethr      Re[znorm]      Re[delta] [meV]
1  1.340834E-01  2.286409E+00  8.742224E-01
2  1.899921E-01  2.286409E+00  8.742224E-01
.....
14  5.485982E-04  2.286409E+00  8.742224E-01
Convergence was reached in nsiter =      14

```

- At the end of the calculation, you should get a few output files at every given temperature. Note that the number of Matsubara frequency points decreases as the temperature increases because fewer frequencies $i\omega_j = i(2n + 1)\pi T$ (n integer) are smaller than the cutoff frequency `wscut`.

The calculation of superconducting properties will be accompanied by significant I/O. In the following we will describe various physical quantities saved in the output files and how to process them. We will use XX in the name of the output files to indicate the temperature at which the equations are solved.

- Eliashberg spectral function and integrated electron-phonon coupling strength (λ).

pb.a2f and pb.a2f_proj files are generated by setting `eliashberg = .true.`

pb.a2f file contains the isotropic Eliashberg spectral function $\alpha^2 F(\omega)$ and cumulative electron-phonon coupling strength λ as a function of frequency ω (meV) for different phonon smearing values (see the end of the file for information about the smearing).

pb.a2f_proj file contains the Eliashberg spectral function as a function of frequency ω (meV), where the 2nd column is the Eliashberg spectral function corresponding to the first smearing in pb.a2f. The remaining ($3 \times$ number of atoms) columns contain the mode-resolved Eliashberg spectral functions corresponding to the first smearing in pb.a2f (there is no specific information on which modes correspond to which atomic species).

- Superconducting gap along the imaginary frequency axis and the real frequency axis.

pb.imag_iso_XX files are generated by setting `eliashberg = .true.`, `liso = .true.`, and `limag = .true.`. Each file contains 3 columns: the Matsubara frequency $i\omega_j$ (eV) along the imaginary axis, the quasiparticle renormalization function $Z(i\omega_j)$, and the superconducting gap $\Delta(i\omega_j)$ (eV).

pb.pade_iso_XX files are generated by setting `lpade = .true.`. Each file contains 5 columns: the frequency ω (eV) along the real axis, the real part of the quasiparticle renormalization function $\text{Re}Z(\omega)$, the imaginary part of the quasiparticle renormalization function $\text{Im}Z(\omega)$, the real part of the superconducting gap $\text{Re}\Delta(\omega)$ (eV), and the imaginary part of the superconducting gap $\text{Im}\Delta(\omega)$ (eV).

pb.acon_iso_XX files are generated by setting `lacon = .true.` and contain similar information as pb.pade_iso_XX.

► 4th step: Plot the superconducting gap along the imaginary and real frequency axis.

You can use the following gnuplot scripts to plot pb.imag_iso_000.30, pb.pade_iso_000.30, and pb.acon_iso_000.30. You should get something similar to Fig. 1 at 0.3 K.

```
$ gnuplot
gnuplot> set xlabel "iw (meV)"
gnuplot> set ylabel "Delta (meV)"
gnuplot> plot "pb.imag_iso_000.30" u ($1*1000):($3*1000) w l lw 2 lt rgb "black" \
> title "Delta-Imag"

gnuplot> set xlabel "w (meV)"
gnuplot> set ylabel "Delta (meV)"
gnuplot> plot "pb.pade_iso_000.30" u ($1*1000):($4*1000) w l lw 1 lt rgb "black" \
> title "Re(Delta)-Pade", \
> "" u ($1*1000):($5*1000) w l lw 1 lt rgb "red" title "Im(Delta)-Pade", \
> "pb.acon_iso_000.30" u ($1*1000):($4*1000) w l lw 1 lt rgb "green" \
> title "Re(Delta)-analytic cont.", \
> "" u ($1*1000):($5*1000) w l lw 1 lt rgb "blue" \
> title "Im(Delta)-analytic cont."
gnuplot> exit
```

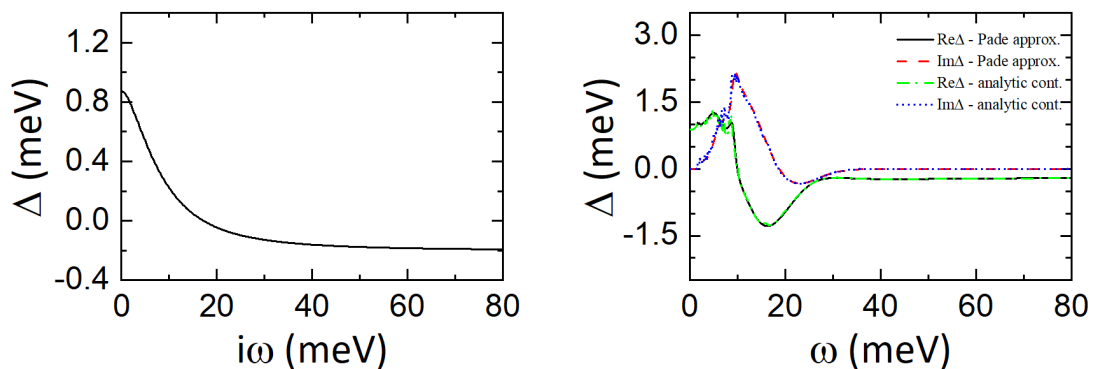



Fig. 1 Left: Superconducting gap along the imaginary axis (columns 1:3 from pb.imag_iso_000.30). Right: Superconducting gap on the real axis (columns 1:4 and 1:5 from pb.pade_iso_000.30 and pb.acon_iso_000.30).

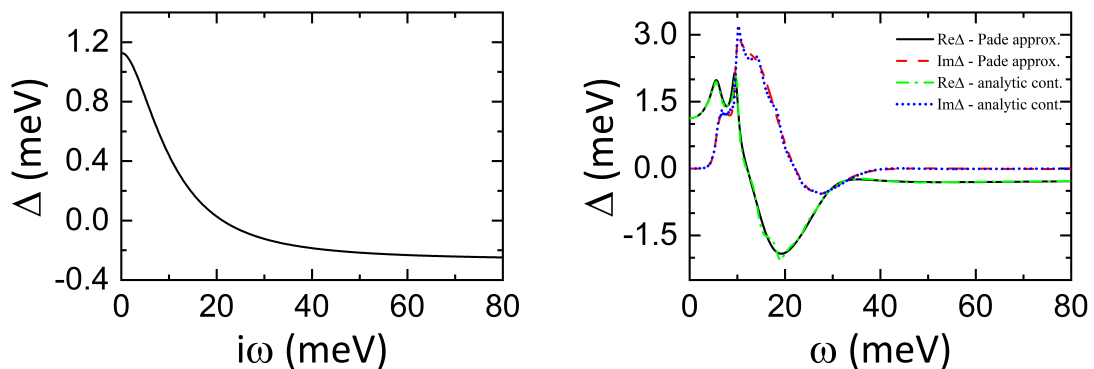


Fig. 2 At convergence you should get something close to this figure (see [Phys. Rev. B 87, 024505 \(2013\)](#) for fully converged calculation parameters).

► 5th step: Plot the leading edge of the superconducting gap as a function of temperature.

Use the shell script `script_gap0_imag` (also shown below) to extract the leading edge of the superconducting gap as a function of temperature and save the information in a new file `pb.imag_iso_gap0`.

```

                                                                    script_gap0_imag
#!/bin/tcsh

awk 'FNR==2 {print FILENAME,$0}' pb.imag_iso_* | awk '{print $1 " " " $4*1000}' > pb.imag_iso_gap0
sed -i 's/pb.imag_iso_//' pb.imag_iso_gap0

```

\$ `./script_gap0_imag`

You can use the following gnuplot script to plot `pb.imag_iso_gap0`. You should get something similar to Fig. 3

```

$ gnuplot
gnuplot> set xlabel "T (K)"
gnuplot> set ylabel "Delta_0 (meV)"
gnuplot> set xrange [0:6]
gnuplot> set yrange [0:1.2]
gnuplot> plot "pb.imag_iso_gap0" with lp ls 3 notitle
gnuplot> exit

```

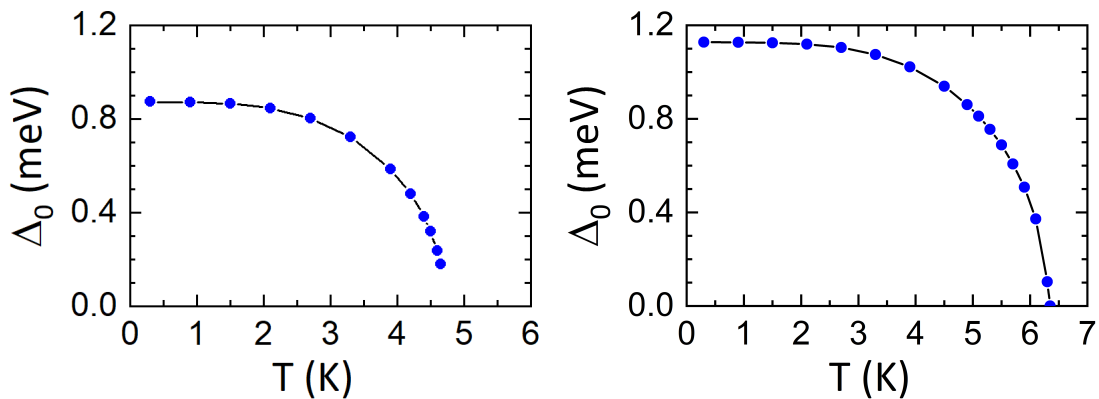


Fig. 3 Calculated isotropic gap of Pb as a function of temperature. At convergence you should get something close to the figure on the right (see [Phys. Rev. B 87, 024505 \(2013\)](#) for fully converged calculation parameters).

You can further extract the leading edge of the superconducting gap as a function of temperature from the calculations on the real axis and compare it with the one obtained on the imaginary axis shown in Fig. 3. You can use the shell scripts `script_gap0_pade` and `script_gap0_acon` to get the `pb.pade_iso_gap0` and `pb.acon_iso_gap0` files. Next plot these files using `gnuplot` as was done above for the `pb.imag_iso_gap0`.

```
$ ./script_gap0_pade
```

```
#!/bin/tcsh                                                                 script_gap0_pade
awk 'FNR==2 {print FILENAME,$0}' pb.pade_iso_* | awk '{print $1 " " " $5*1000}' > pb.pade_iso_gap0
sed -i 's/pb.pade_iso_//' pb.pade_iso_gap0
```

and

```
$ ./script_gap0_acon
```

```
#!/bin/tcsh                                                                 script_gap0_acon
awk 'FNR==2 {print FILENAME,$0}' pb.acon_iso_* | awk '{print $1 " " " $5*1000}' > pb.acon_iso_gap0
sed -i 's/pb.acon_iso_//' pb.acon_iso_gap0
```

► 6th step: Solve the linearized isotropic Migdal-Eliashberg equation for the critical temperature.

Near T_c , $\Delta(i\omega_j) \rightarrow 0$ and the set of Eqs. (1) reduces to a linear matrix equation for $\Delta(i\omega_j)$:

$$\Delta(i\omega_j) = \sum_{j'} K_{jj'} \Delta(i\omega_{j'}) \quad (3)$$

where

$$K_{jj'} = \frac{1}{|2j' + 1|} \left[\lambda(\omega_j - \omega_{j'}) - \mu_c^* - \delta_{jj'} \sum_{j''} \lambda(\omega_j - \omega_{j''}) s_j s_{j''} \right] \quad (4)$$

and $s_j = \text{sign}(\omega_j)$. The critical temperature T_c is defined as the value at which the maximum eigenvalue of $K_{jj'}$ is 1.

This step can be done by starting from a file containing the Eliashberg spectral function (pb.a2f_iso) using the following jobscript (job.epw2) and input file (epw2.in; only differences with respect to epw1.in file are shown below):

Note 1: In this case ephmatXX, freq, egnv, and ikmap files (saved in the prefix.ephmat directory) are not used. You can also solve the isotropic Migdal-Eliashberg equations at other temperatures starting from a file containing the Eliashberg spectral function (pb.a2f_iso). This procedure does not work for solving the anisotropic Migdal-Eliashberg equations.

Note 2: You only need to use one CPU if the isotropic Migdal-Eliashberg or the linearized Migdal-Eliashberg equation are solved starting from the Eliashberg spectral function.

\$ sbatch job.epw2

```

                                                                    job.epw2
#!/bin/bash
#SBATCH -J job.epw2           # Job name
#SBATCH -N 1                 # Total # of nodes
#SBATCH --ntasks-per-node 1
#SBATCH -t 01:00:00         # Run time (hh:mm:ss)
#SBATCH -A EPSchool2022
#SBATCH -p small
#SBATCH --reservation=EPSchoolDay3

# Launch MPI code...
export PATHQE=/work2/06868/giustino/EP-SCHOOL/q-e

ibrun $PATHQE/bin/epw.x -nk 1 -in epw2.in > epw2.out

```

```

--                                                                    epw2.in
ep_coupling = .false.
elph        = .false.
epwwrite    = .false.
epwread     = .true.
wannierize  = .false.
ephwrite    = .false.

fila2f      = 'pb.a2f'
lpade       = .false.
lacon       = .false.
tc_linear   = .true.           ! solve linearized ME eqn. for Tc
tc_linear_solver = 'power'    ! algorithm to solve Tc eigenvalue problem: 'power' OR 'lapack'

nstamp      = 21               ! number of temperature points
temps       = 0.25 5.25       ! evenly spaced nstamp temperature points according to
                               ! (temps(2)-temps(1))/(nstamp-1).

```

Start: Solving (isotropic) linearized Eliashberg equation with solver = power

For the first Temp. 0.25 K
 Total number of frequency points nsiw(1) = 739
 Cutoff frequency wscut = 0.1001

Superconducting transition temp. Tc is the one which has Max. eigenvalue close to 1

Temp. (K)	Max. eigenvalue	nsiw (itemp)	wscut (eV)	Nr. of iters to Converge
0.25	4.5835907	739	0.1001	6
0.50	3.6708046	369	0.1000	7

```
.....          .....          ...          .....          ..
5.25          0.8996285          35          0.1009          27
```

Finish: Solving (isotropic) linearized Eliashberg equation

You can extract the maximum eigenvalue as a function of temperature from the epw2.out using script_max_eigenvalue and save the date in data_max_eigenvalue.dat file.

```
$ ./script_max_eigenvalue
```

```
script_max_eigenvalue
#!/bin/tcsh
grep -A 25 "Max. eigenvalue" epw2.out | tail -21 | awk '{print $1 " " " $2}' > data_max_eigenvalue.dat
```

Plot data_max_eigenvalue.dat to obtain the T_c . The critical temperature is defined as the value for which the maximum eigenvalue is close to 1. You can use the gnuplot script below to get the graph shown in Fig. 4

```
$ gnuplot
gnuplot> set xlabel "T (K)"
gnuplot> set ylabel "Max. eigenvalue"
gnuplot> set xrange [0:6]
gnuplot> set arrow from 0,1 to 6,1 nohead lt 2 lw 1
gnuplot> plot "data_max_eigenvalue.dat" with lp ls 3 notitle
gnuplot> exit
```

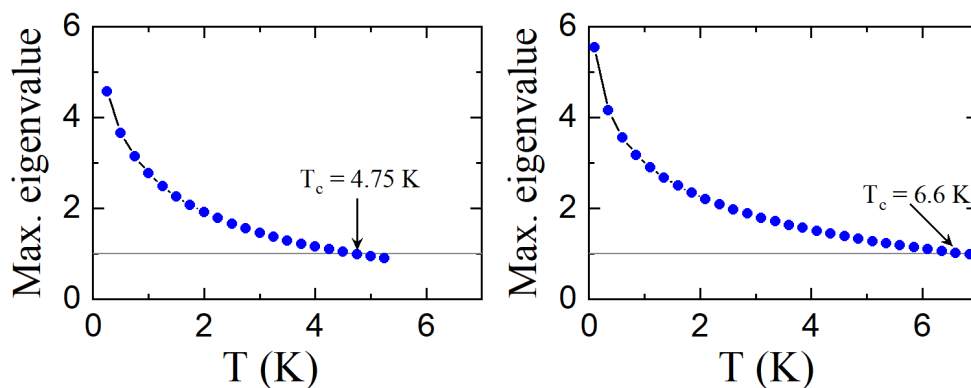


Fig. 4 Calculated maximum eigenvalue as a function of temperature. At convergence you should get something closer to the right hand-side figure.

Exercise 2

In this tutorial we are going to calculate the superconducting properties of MgB₂ by solving the anisotropic Migdal-Eliashberg equations. The theory related to this tutorial can be found in the [Phys. Rev. B 87, 024505 \(2013\)](#).

Go to exercise2:

```
$ cd ../../exercise2
```

► 1st step: Run a self-consistent calculation on a homogeneous 12×12×12 **k**-point grid and a phonon calculation on a homogeneous 3×3×3 **q**-point grid using the following jobscript (`job.ph`) and input files (`scf.in` and `ph.in`) for MgB₂:

Note: The smearing is quite large in order to get reasonable values in subsequent calculations.

```
$ cd phonon
$ sbatch job.ph
```

```
#!/bin/bash                                                                    job.ph
#SBATCH -J job.ph                        # Job name
#SBATCH -N 1                              # Total # of nodes
#SBATCH --ntasks-per-node 56
#SBATCH -t 01:00:00                      # Run time (hh:mm:ss)
#SBATCH -A EPSchool2022
#SBATCH -p small
#SBATCH --reservation=EPSchoolDay3

# Launch MPI code...
export PATHQE=/work2/06868/giustino/EP-SCHOOL/q-e

ibrun $PATHQE/bin/pw.x -nk 56 -in scf.in > scf.out
ibrun $PATHQE/bin/ph.x -nk 56 -in ph.in > ph.out
```

```
--                                                                              scf.in
&control
  calculation = 'scf'
  restart_mode = 'from_scratch',
  prefix      = 'mgb2',
  pseudo_dir  = '../..pseudo/',
  outdir      = './',
/
&system
 ibrav      = 4,
  celldm(1) = 5.8260252227888,
  celldm(3) = 1.1420694129095,
  nat       = 3,
  ntyp      = 2,
  ecutwfc   = 40
  smearing  = 'mp'
  occupations = 'smearing'
  degauss   = 0.05
/
&electrons
  diagonalization = 'david'
  mixing_mode     = 'plain'
  mixing_beta     = 0.7
  conv_thr        = 1.0d-9
/
ATOMIC_SPECIES
Mg 24.305 Mg.pz-n-vbc.UPF
B  10.811 B.pz-vbc.UPF
ATOMIC_POSITIONS crystal
Mg 0.000000000 0.000000000 0.000000000
B  0.333333333 0.666666667 0.500000000
B  0.666666667 0.333333333 0.500000000
K_POINTS AUTOMATIC
12 12 12 0 0 0
```

```
--                                                                              ph.in
&inputph
  prefix = 'mgb2',
  fildyn = 'mgb2.dyn.xml',
  tr2_ph = 1.0d-16
  fildvscf = 'dvscf',
```

```
ldisp = .true.,
nq1 = 3,
nq2 = 3,
nq3 = 3
```

Dynamical matrices for (3, 3, 3) uniform grid of q-points
(6 q-points):

N	xq(1)	xq(2)	xq(3)
1	0.000000000	0.000000000	0.000000000
2	0.000000000	0.000000000	0.291867841
3	0.000000000	0.384900179	0.000000000
4	0.000000000	0.384900179	0.291867841
5	0.333333333	0.577350269	0.000000000
6	0.333333333	0.577350269	0.291867841

► 2nd step: Gather the .dyn, .dvscf, and patterns files into a new save directory using the pp.py python script.

```
$ python3 /work2/06868/giustino/EP-SCHOOL/q-e/EPW/bin/pp.py
```

The script will ask you to provide the prefix of your calculation (here "mgb2").

► 3rd step: Do a non self-consistent calculation on a $6 \times 6 \times 6$ **uniform** and Γ -**centered grid between [0,1[in crystal coordinates** and an EPW calculation for the anisotropic superconducting properties using the following jobscript (job.epw1) and input files (nscf.in and epw1.in):

Note 1: The homogeneous k grid for the non self-consistent calculations can be generated using the script [kmesh.pl](#)

```
$ /work2/06868/giustino/EP-SCHOOL/q-e/wannier90-3.1.0/utility/kmesh.pl 6 6 6
```

Note 2: A non self-consistent calculation requires the charge density found from a previous self-consistent run with [pw.x](#). In the jobscript [job.epw1](#) you can see that a self-consistent calculation is run first with the same [scf.in](#) file used in the [phonon](#) directory. Alternatively, one can make the [mgb2.save](#) directory and copy there the files from [phonon/mgb2.save](#). For this in [job.epw1](#), you need to comment the line

```
#ibrun $PATHQE/bin/pw.x -nk 8 -in scf.in > scf.out
```

and uncomment the following three lines

```
mkdir mgb2.save
cp ../phonon/mgb2.save/charge-density.dat mgb2.save/
cp ../phonon/mgb2.save/data-file-schema.xml mgb2.save/
```

Note 3: EPW calculations with [ephwrite = .true.](#) require that the fine k or q grids are commensurate, i.e., [nkf1](#), [nkf2](#), [nkf3](#) to be multiple of [nqf1](#), [nqf2](#), [nqf3](#).

Note 4: The Migdal-Eliashberg equations are solved in the standard, Fermi surface restriction (FSR), approximation.

```
$ cd ../epw1-FSR
$ sbatch job.epw1
```

```
#!/bin/bash                                                                                                     job.epw1
#SBATCH -J job.epw1          # Job name
#SBATCH -N 1                 # Total # of nodes
#SBATCH --ntasks-per-node 56
#SBATCH -t 01:00:00         # Run time (hh:mm:ss)
#SBATCH -A EPSchool2022
#SBATCH -p small
#SBATCH --reservation=EPSchoolDay3
```

```

# Launch MPI code...
export PATHQE=/work2/06868/giustino/EP-SCHOOL/q-e

ibrun $PATHQE/bin/pw.x -nk 56 -in scf.in > scf.out
#alternatively to re-run a scf calculation copy files from ../phonon/mgb2.save
#mkdir mgb2.save
#cp ../phonon/mgb2.save/charge-density.dat mgb2.save/
#cp ../phonon/mgb2.save/data-file-schema.xml mgb2.save/

ibrun $PATHQE/bin/pw.x -nk 56 -in nscf.in > nscf.out
ibrun $PATHQE/bin/epw.x -nk 56 -in epw1.in > epw1.out

```

```

&control                                                                    nscf.in
  calculation = 'nscf'
  prefix      = 'mgb2',
  pseudo_dir  = '../pseudo/',
  outdir      = './',
/
&system
  ibrav       = 4,
  cellldm(1)  = 5.8260252227888,
  cellldm(3)  = 1.1420694129095,
  nat         = 3,
  ntyp        = 2,
  ecutwfc     = 40
  smearing    = 'mp'
  occupations = 'smearing'
  degauss     = 0.05
/
&electrons
  diagonalization = 'david'
  mixing_mode     = 'plain'
  mixing_beta     = 0.7
  conv_thr        = 1.0d-9
/
ATOMIC_SPECIES
Mg 24.305 Mg.pz-n-vbc.UPF
B  10.811 B.pz-vbc.UPF
ATOMIC_POSITIONS crystal
Mg 0.000000000 0.000000000 0.000000000
B  0.333333333 0.666666667 0.500000000
B  0.666666667 0.333333333 0.500000000
K_POINTS crystal
216
0.00000000 0.00000000 0.00000000 4.629630e-03
0.00000000 0.00000000 0.16666667 4.629630e-03
...

```

```

--                                                                    epw1.in
&inputepw
  prefix      = 'mgb2',
  outdir      = './'
  dvscf_dir   = '../phonon/save' ! directory where .dyn, .dvscf and prefix.phsave/patterns.xx.yy
                                ! files obtained from phonon calculation are stored

  ep_coupling = .true.           ! run e-ph coupling calculation
  elph        = .true.           ! calculate e-ph coefficients
  epwwrite    = .true.           ! write e-ph matrices in the Wann representation
  epwread     = .false.          ! read e-ph matrices from the 'prefix.epmatwp' file

  etf_mem     = 1                ! more IO (slower) but less memory is required

  wannierize  = .true.           ! calculate Wannier functions using W90 library
  nbndsub     = 5                ! number of Wannier functions to utilize

  num_iter    = 500
  dis_froz_max = 8.8
  proj(1)     = 'B:pz'
  proj(2)     = 'f=0.5,1.0,0.5:s'

```

```

proj(3)      = 'f=0.0,0.5,0.5:s'
proj(4)      = 'f=0.5,0.5,0.5:s'

iverbosity   = 2           ! 2 = verbose output for the SC part

fsthick      = 0.2         ! Fermi window thickness [eV]
degaussw     = 0.05       ! smearing in the energy-conserving delta functions in [eV]

fermi_plot   = .true.     ! write files to plot Fermi surface
ephwrite     = .true.     ! write ephmatXX, egnv, freq, and ikmap files in 'prefix.ephmat' directory
eliashberg   = .true.     ! calculate Eliashberg spectral function

laniso       = .true.     ! solve anisotropic ME eqs.
limag        = .true.     ! solve ME eqs. imaginary axis
lpade        = .true.     ! solve ME eqs. on real axis using Pade approximants

nsiter       = 500        ! number of self-consistent iterations when solving ME eqs.
conv_thr_imag = 1.0d-3    ! convergence threshold for solving ME eqs. on imaginary axis
wscut        = 0.5        ! upper limit over Matsubara freq. summation in ME eqs on imag. axis [eV]
muc          = 0.05       ! effective Coulomb potential used in ME eqs.

nstemp       = 3          ! number of temperature points at which the ME eqs. are solved
temps        = 10 20      ! even space mode: step between points is (temps(2)-temps(1))/(nstemp-1)

nk1          = 6          ! dimensions of the coarse electronic grid
nk2          = 6
nk3          = 6

nq1          = 3          ! dimensions of the coarse phonon grid
nq2          = 3
nq3          = 3

mp_mesh_k    = .true.    ! use irreducible electronic fine mesh
nkf1         = 40
nkf2         = 40         ! dimensions of the fine electronic grid
nkf3         = 40

nqf1         = 20
nqf2         = 20         ! dimensions of the fine phonon grid
nqf3         = 20
/

```

With the above input, we are instructing EPW to:

- Fourier-transform the electron-phonon matrix elements from a coarse $6 \times 6 \times 6$ to a dense $40 \times 40 \times 40$ \mathbf{k} -point grid and from a coarse $3 \times 3 \times 3$ to a dense $20 \times 20 \times 20$ \mathbf{q} -point grid.

```

Using uniform q-mesh:  20 20 20
Size of q point mesh for interpolation:      8000
Using uniform MP k-mesh:  40 40 40
Size of k point mesh for interpolation:      6468
Max number of k points per pool:           116

```

- Pre-compute the \mathbf{q} -points that fall within the `fsthick` window. If at a specific \mathbf{q} -point at least one $\mathbf{k} + \mathbf{q}$ eigenvalue falls within the user-defined `fsthick`, then the \mathbf{q} -point is selected.

```

Number selected, total      100          100
Number selected, total      200          204
.....
Number selected, total      7800         7954
We only need to compute     7846 q-points

```


- Write on disk in the `mgb2.ephmat` directory the: (1) `ephmatXX` files (one per CPU) containing the electron-phonon matrix elements within the Fermi window (`fsthick`) on the dense `k` and `q` grids, (2) `freq` file containing the phonon frequencies on the dense `q` grid, (3) `egnv` file containing the eigenvalues within the Fermi window on the dense `k` grid, and (4) `ikmap` file containing the index of the `k`-points on the dense (irreducible) grid within the Fermi window. All these files are produced by setting `ephwrite = .true.`. These files are unformatted and required for solving the anisotropic Migdal-Eliashberg equations.

```
Nr. of irreducible k-points on the uniform grid:      3234

Finish mapping k+sign*q onto the fine irreducibe k-mesh and writing .ikmap file

Nr irreducible k-points within the Fermi shell =      446 out of      3234

Progression iq (fine) =      100/      7846
Progression iq (fine) =      200/      7846
.....
.....
Progression iq (fine) =      7800/      7846
      Fermi level (eV) =      0.746936938272991D+01
DOS(states/spin/eV/Unit Cell) =      0.324589885940489D+00
      Electron smearing (eV) =      0.500000000000000D-01
      Fermi window (eV) =      0.200000000000000D+00

Finish writing .ephmat files
```

- Write the Fermi surface files `mgb2.fs_YY.cube` (`YY` = band index within the `fsthick`) and `mgb2.fs.frmsf` by setting `fermi_plot = .true.`. The `*.cube` files can be visualized with `VESTA` and the `*.frmsf` file can be visualized with `FermiSurfer`.

```
Fermi surface calculation on fine mesh
      Fermi level (eV) =      7.469369
      3  bands within the Fermi window
```

- Calculate the isotropic and anisotropic electron-phonon coupling strength by setting the keywords `eliashberg = .true.` in the EPW input file.

The anisotropic electron-phonon coupling strength takes the following form:

$$\lambda_{n\mathbf{k},m\mathbf{k}+\mathbf{q}}(\omega_j) = N_F \sum_{\nu} \frac{2\omega_{\mathbf{q}\nu}}{\omega_j^2 + \omega_{\mathbf{q}\nu}^2} |g_{m\nu}(\mathbf{k}, \mathbf{q})|^2 \quad (5)$$

The band- and wavevector-dependent electron-phonon coupling strength $\lambda_{n\mathbf{k}}(\omega_j)$ is defined as:

$$\lambda_{n\mathbf{k}}(\omega_j) = \sum_m \int \frac{d\mathbf{q}}{\Omega_{\text{BZ}}} \frac{\delta(\epsilon_{m\mathbf{k}+\mathbf{q}} - \epsilon_F)}{N_F} \lambda_{n\mathbf{k},m\mathbf{k}+\mathbf{q}}(\omega_j) \quad (6)$$

The isotropic electron-phonon coupling strength takes the form:

$$\lambda(\omega_j) = \sum_n \int \frac{d\mathbf{k}}{\Omega_{\text{BZ}}} \frac{\delta(\epsilon_{n\mathbf{k}} - \epsilon_F)}{N_F} \lambda_{n\mathbf{k}}(\omega_j) \quad (7)$$

The standard electron-phonon coupling strength λ found in the literature corresponds to setting $\omega_j = 0$ in Eq. (7).

The isotropic Eliashberg spectral function takes the following form:

$$\alpha^2 F(\omega) = \frac{1}{N_F} \sum_{nm\nu} \int \frac{d\mathbf{k}}{\Omega_{\text{BZ}}} \int \frac{d\mathbf{q}}{\Omega_{\text{BZ}}} |g_{mn\nu}(\mathbf{k}, \mathbf{q})|^2 \delta(\omega - \omega_{\mathbf{q}\nu}) \delta(\epsilon_{n\mathbf{k}} - \epsilon_F) \delta(\epsilon_{m\mathbf{k}+\mathbf{q}} - \epsilon_F) \quad (8)$$

- Solve the anisotropic FSR Migdal-Eliashberg equations on the imaginary frequency axis by setting the keywords `eliashberg = .true.`, `laniso = .true.`, and `limag = .true.` in the EPW input file. The equations are solved self-consistently for each temperature value specified in the input file. The calculation at each temperature ends when either the converge threshold (`conv_thr_iaxis`) or the maximum number of iterations (`nsiter`) is reached.

Note 1: If at a specific temperature the maximum number of iterations is reached without achieving convergence, the code will stop and not move to the next temperature in the list.

Note 2: Because the electron-phonon matrix elements do not depend on the temperature at which the Migdal-Eliashberg equations are solved, they can be reused in subsequent EPW calculations at different temperatures. This is the reason why `ephmatXX` files are saved in the `mgb2.ephmat` directory.

The anisotropic FSR Migdal-Eliashberg equations take the following form:

$$Z_{n\mathbf{k}}(i\omega_j) = 1 + \frac{\pi T}{\omega_j N_F} \sum_{mj'} \int \frac{d\mathbf{q}}{\Omega_{\text{BZ}}} \frac{\omega_{j'}}{\sqrt{\omega_{j'}^2 + \Delta_{m\mathbf{k}+\mathbf{q}}^2(i\omega_{j'})}} \times \lambda_{n\mathbf{k}, m\mathbf{k}+\mathbf{q}}(\omega_j - \omega_{j'}) \delta(\epsilon_{m\mathbf{k}+\mathbf{q}} - \epsilon_F)$$

$$Z_{n\mathbf{k}}(i\omega_j) \Delta_{n\mathbf{k}}(i\omega_j) = \frac{\pi T}{N_F} \sum_{mj'} \int \frac{d\mathbf{q}}{\Omega_{\text{BZ}}} \frac{\Delta_{m\mathbf{k}+\mathbf{q}}(i\omega_{j'})}{\sqrt{\omega_{j'}^2 + \Delta_{m\mathbf{k}+\mathbf{q}}^2(i\omega_{j'})}} \times [\lambda_{n\mathbf{k}, m\mathbf{k}+\mathbf{q}}(\omega_j - \omega_{j'}) - \mu_c^*] \delta(\epsilon_{m\mathbf{k}+\mathbf{q}} - \epsilon_F), \quad (9)$$

where $\lambda_{n\mathbf{k}, m\mathbf{k}+\mathbf{q}}(\omega_j - \omega_{j'})$ is the anisotropic electron-phonon coupling strength. The semi-empirical Coulomb parameter μ_c^* is provided as an input variable `muc` in the EPW calculation.

```
=====
Solve anisotropic Eliashberg equations
=====
.....
Electron-phonon coupling strength =      0.7115986

Estimated Allen-Dynes Tc =      36.325641 K for muc =      0.05000

Estimated w_log in Allen-Dynes Tc =      59.079768 meV

Estimated BCS superconducting gap =      5.357666 meV

Estimated Tc from machine learning model =      37.738740 K

temp( 1) =      10.00000 K

Solve anisotropic Eliashberg equations on imaginary-axis

Total number of frequency points nsiw( 1) =      92
```

```

Cutoff frequency wscut =      0.5008

Size of allocated memory per pool: ~=      0.0664 Gb
iter      ethr      znormi      deltai [meV]
  1  2.676078E+00  1.669495E+00  6.099423E+00
  2  7.355352E-02  1.667331E+00  6.348325E+00
  .....
 14  6.875986E-04  1.662053E+00  6.949151E+00
Convergence was reached in nsiter =      14

Chemical potential (itemp =  1) =      7.4693693827E+00 eV

Temp (itemp =  1) =  10.000 K Free energy =      -0.006853 meV

Min. / Max. values of superconducting gap =      0.000000  12.286016 meV

```

- Perform the analytic continuation of the solutions along the imaginary frequency axis to the real frequency axis by using Padé approximants (`lpade = .true.`). Note the analytic continuation with the iterative procedure (`lacon = .true.`) is not performed since this is very expensive computationally in the anisotropic case (hours to days).

```

Padé approximant of anisotropic Eliashberg equations from imaginary-axis to real-axis
Cutoff frequency wscut =      0.5000

pade Re[znorm] [eV] Re[delta] [eV]
82  1.691852E+00  6.409036E+00

Convergence was reached for N =      82 Padé approximants

```

The calculation of superconducting properties will be accompanied by significant I/O. In the following we will describe various physical quantities saved in the output files and how to process them. We will use XX in the name of the output files to indicate the temperature at which the equations are solved.

► 4th step: Plot the isotropic and anisotropic electron-phonon coupling strength.

`mgb2.lambda_pairs`, `mgb2.lambda_k_pairs`, and `mgb2.a2f` files are generated by setting `eliashberg = .true.`

`mgb2.lambda_pairs` file contains the anisotropic electron-phonon coupling strength $\lambda_{n\mathbf{q},m\mathbf{k}+\mathbf{q}}(0)$ on the Fermi surface.

`mgb2.lambda_k_pairs` file contains the band- and wavevector-dependent anisotropic electron-phonon coupling strength $\lambda_{n\mathbf{k}}(0)$ on the Fermi surface.

`mgb2.a2f` file contains the isotropic Eliashberg spectral function $\alpha^2F(\omega)$ and cumulative electron-phonon coupling strength as a function of frequency ω (meV) for different phonon smearing values (see the end of the file for information about the smearing).

Note: First, put `#` in front of the 1st line and the last 7 lines of `mgb2.a2f`, otherwise gnuplot does not work.

You can use the following gnuplot scripts to plot. You should get something similar to Fig. 5.

```

$ gnuplot
gnuplot> set xlabel "lambda_(nk,mk+q)(0)"
gnuplot> set xrange [0:4]
gnuplot> set yrange [0:1.1]
gnuplot> plot "mgb2.lambda_pairs" w l lw 2 lt rgb "black" notitle

gnuplot> set xlabel "lambda_(nk)(0)"
gnuplot> set xrange [0:1.5]
gnuplot> set yrange [0:1.1]
gnuplot> plot "mgb2.lambda_k_pairs" w l lw 2 lt rgb "black" notitle

gnuplot> set xlabel "w (meV)"
gnuplot> set xrange [0:110]
gnuplot> set yrange [0:1.5]
gnuplot> plot "mgb2.a2f" w l lw 2 lt rgb "black" title "a2f", \
> "" u ($1):($12) w l lw 2 lt rgb "red" title "lambda"
gnuplot> exit

```

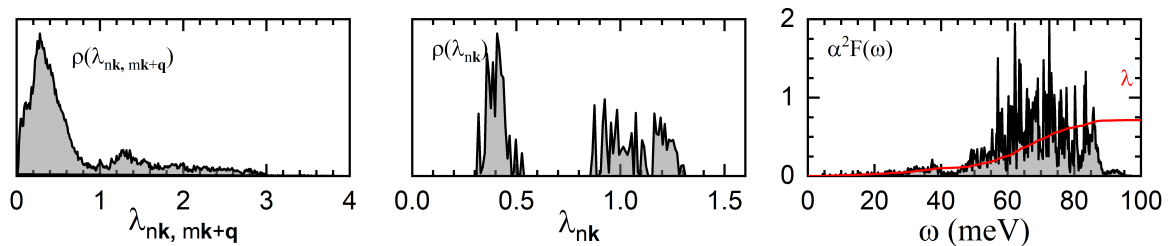


Fig. 5 Left: The anisotropic electron-phonon coupling strength $\lambda_{nq,mk+q}(0)$ (from `mgb2.lambda_pairs`). Middle: The anisotropic electron-phonon coupling strength $\lambda_{nk}(0)$ on the Fermi surface (from `mgb2.lambda_k_pairs`). Right: The isotropic Eliashberg spectral function $\alpha^2 F(\omega)$ (columns 1:2 from `mgb2.a2f`) and integrated electron-phonon coupling strength λ (columns 1:12 from `mgb2.a2f`).

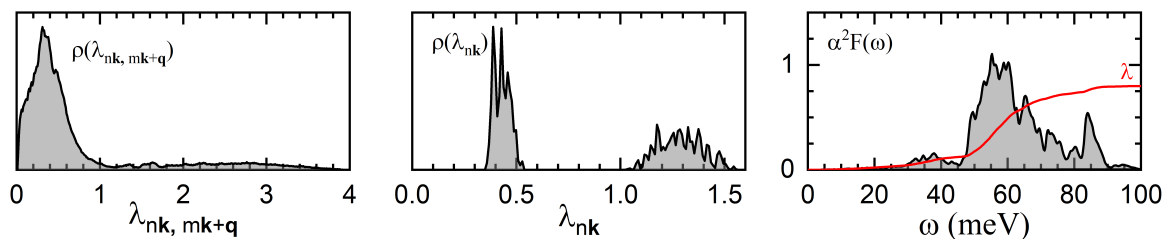


Fig. 6 At convergence you should get something close to this figure (see [Phys. Rev. B 87, 024505 \(2013\)](#) for fully converged calculation parameters).

► 5th step: Plot the superconducting gap along the imaginary frequency axis and the real frequency axis.

`mgb2.imag_aniso_XX` files are generated by setting `eliashberg = .true.`, `laniso = .true.`, and `limag = .true.`. Each file contains 4 columns: the frequency $i\omega_j$ (eV) along the imag-

inary axis, the Kohn-Sham eigenvalue ϵ_{nk} (eV) relative to the Fermi level, the quasiparticle renormalization $Z_{nk}(i\omega_j)$, and the superconducting gap $\Delta_{nk}(i\omega_j)$ (eV).

mgb2.pade_aniso_XX files are generated by setting `lpade = .true.`. Each file contains 6 columns: the energy ω (eV) along the real axis, the Kohn-Sham eigenvalue ϵ_{nk} (eV) relative to the Fermi level, the real part of the quasiparticle renormalization $\text{Re}Z_{nk}(\omega)$, the imaginary part of the quasiparticle renormalization $\text{Im}Z_{nk}(\omega)$, the real part of the superconducting gap $\text{Re}\Delta_{nk}(\omega)$ (eV), and the imaginary part of the superconducting gap $\text{Im}\Delta_{nk}(\omega)$ (eV).

mgb2.acon_aniso_XX files could also be generated by setting `lacon = .true.`. These files will contain similar information as mgb2.pade_aniso_XX.

You can use the following gnuplot scripts to plot. You should get something similar to Fig. 7 at 10 K.

```
$ gnuplot
gnuplot> set xlabel "iw (meV)"
gnuplot> set ylabel "Delta_nk (meV)"
gnuplot> set xrange [0:180]
gnuplot> set yrange [0:15]
gnuplot> plot "mgb2.imag_aniso_010.00" u ($1*1000):($4*1000) with points pt 7 \
> notitle

gnuplot> set xlabel "w (meV)"
gnuplot> set ylabel "Delta_nk (meV)"
gnuplot> set xrange [0:180]
gnuplot> set yrange [-20:45]
gnuplot> plot "mgb2.pade_aniso_010.00" u ($1*1000):($5*1000) with points pt 7 \
> notitle
gnuplot> exit
```

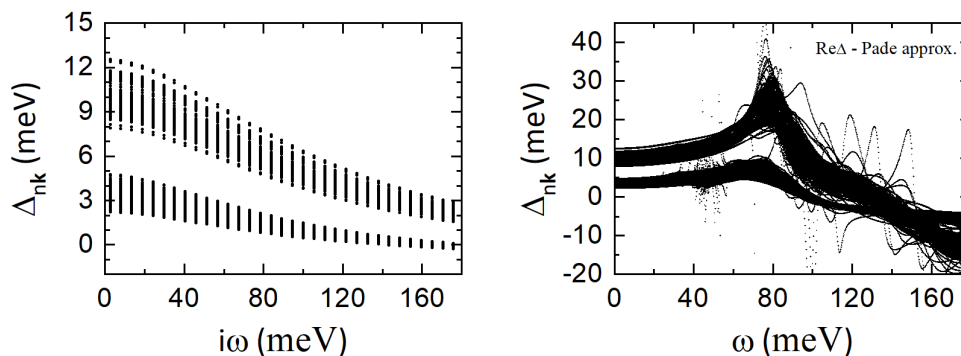


Fig. 7 Left: Superconducting gap along the imaginary axis (columns 1:4 from mgb2.imag_aniso_010.00). Right: Superconducting gap along the real axis (columns 1:5 from mgb2.pade_aniso_010.00 - this file is about 70MB).

The fine k and q point grids need to be much denser for real calculations. At convergence you should get:

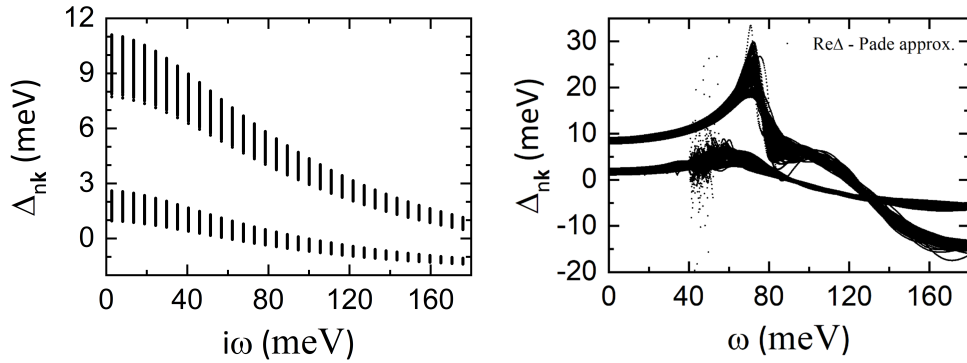


Fig. 8 At convergence you should get something close to this figure (see [Phys. Rev. B 87, 024505 \(2013\)](#) for fully converged calculation parameters). (Note: Only about half of the points are shown.)

► 6th step: Plot the leading edge of the superconducting gap as a function of temperature.

You should get the following graph by plotting the data from all `mgb2.imag_aniso_gap0_XX` files. Use the following gnuplot script.

```
$ gnuplot
gnuplot> set xlabel "T (K)"
gnuplot> set ylabel "Delta_nk (meV)"
gnuplot> set xrange [0:60]
gnuplot> set yrange [0:15]
gnuplot> plot "mgb2.imag_aniso_gap0_010.00" w l lw 1 lt rgb "black" notitle, \
> "mgb2.imag_aniso_gap0_015.00" w l lw 1 lt rgb "black" notitle, \
> "mgb2.imag_aniso_gap0_020.00" w l lw 1 lt rgb "black" notitle, \
> "mgb2.imag_aniso_gap0_025.00" w l lw 1 lt rgb "black" notitle, \
> "mgb2.imag_aniso_gap0_030.00" w l lw 1 lt rgb "black" notitle, \
> "mgb2.imag_aniso_gap0_035.00" w l lw 1 lt rgb "black" notitle, \
> "mgb2.imag_aniso_gap0_040.00" w l lw 1 lt rgb "black" notitle, \
> "mgb2.imag_aniso_gap0_045.00" w l lw 1 lt rgb "black" notitle, \
> "mgb2.imag_aniso_gap0_050.00" w l lw 1 lt rgb "black" notitle, \
> "mgb2.imag_aniso_gap0_055.00" w l lw 1 lt rgb "black" notitle
gnuplot> exit
```

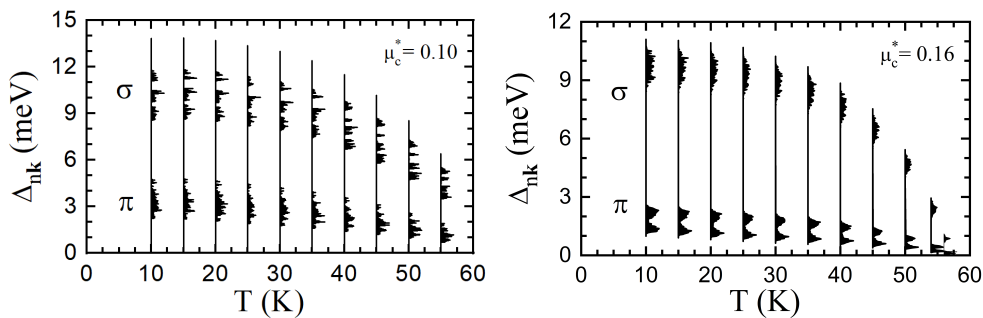


Fig. 9 Calculated anisotropic superconducting gap of MgB_2 on the Fermi surface as a function of temperature. At convergence you should get the right hand-side figure (see [Phys. Rev. B 87, 024505 \(2013\)](#) for fully converged calculation parameters). (Note: the heights of the histograms are multiplied by a factor of 2 while plotting for visibility.)

► 7th step: Plot the superconducting quasiparticle density of states.

The quasiparticle density of states (DOS) in the superconducting state relative to the DOS in the normal state is given by:

$$\frac{N_S(\omega)}{N_F} = \sum_n \int_{\Omega_{BZ}} \frac{d\mathbf{k}}{\Omega_{BZ}} \frac{\delta(\epsilon_{n\mathbf{k}} - \epsilon_F)}{N_F} \text{Re} \left[\omega / \sqrt{\omega^2 - \Delta_{n\mathbf{k}}^2(\omega)} \right] \quad (10)$$

mgb2.qdos_XX files contain the quasiparticle density of states in the superconducting state relative to the density of states in the normal state $N_S(\omega)/N_F$ as a function of frequency (eV) at various XX temperatures.

Use the following gnuplot script to plot mgb2.qdos_010.00. You should get something similar to Fig. 10 (left) at 10 K:

```
$ gnuplot
gnuplot> set xlabel "w (meV)"
gnuplot> set ylabel "N_s(w)/N_F"
gnuplot> set xrange [0:15]
gnuplot> set yrange [0:2.25]
gnuplot> plot "mgb2.qdos_010.00" u ($1*1000):($2/0.3330227) w l lw 2 lt \
> rgb "black" notitle
gnuplot> exit
```

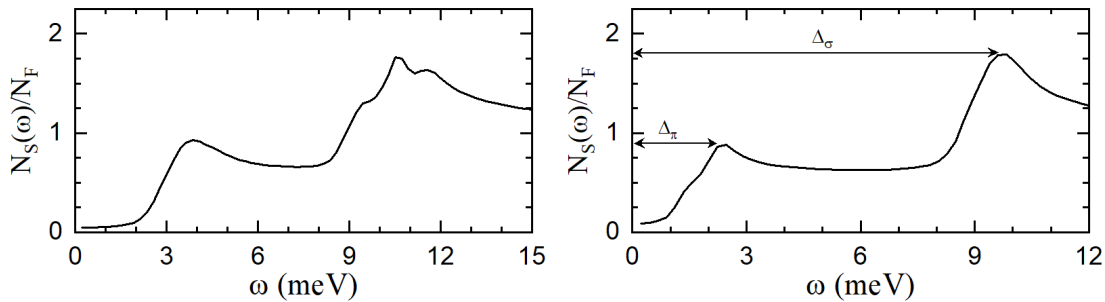


Fig. 10 Calculated $N_S(\omega)/N_F$ as a function of frequency at 10 K. At convergence you should get something closer to the right hand-side figure. (Note: the second column of mgb2.qdos_XX should be divided by the value of DOS from the epw1.out).

► 8th step: (Optional due to time limit) Try to increase the fine grids and see if you can get a result closer to convergence. Note that if either \mathbf{k} or \mathbf{q} is changed you need to obtain new ephmatXX, egnv, freq, and ikmap files (saved in the mgb2.ephmat directory).

► 9th step: (Optional due to time limit) Check the effect of the Coulomb pseudopotential μ_c^* on the superconducting gap and the critical temperature by varying the input variable muc. For this step you can re-use the files saved in the mgb2.ephmat directory.

► 10th step: Solve the anisotropic full-bandwidth (FBW) Migdal-Eliashberg equations (**new feature in EPW**). The self-consistent and non self-consistent calculations are the same as for the standard FSR approximation, you can either copy the './epw1-FSR/mgb2.save' directory or re-run the self-consistent and non self-consistent calculations. After this, do an EPW calculation using the following jobscript (job.epw2) and input file (epw2.in; only differences with respect to './epw1-FSR/epw1.in' file are shown below):

Note: Here, we have fixed the chemical potential at the Fermi level. If you want to update the chemical potential at every temperature, set `muchem = .true.` in your EPW input file.

```
$ cd ../epw2-FBW
$ sbatch job.epw2
```

```

                                                                    job.epw2
#!/bin/bash
#SBATCH -J job.epw2          # Job name
#SBATCH -N 1                 # Total # of nodes
#SBATCH --ntasks-per-node 56
#SBATCH -t 01:00:00         # Run time (hh:mm:ss)
#SBATCH -A EPSchool2022
#SBATCH -p small
#SBATCH --reservation=EPSchoolDay3

# Launch MPI code...
export PATHQE=/work2/06868/giustino/EP-SCHOOL/q-e

ibrun $PATHQE/bin/pw.x -nk 56 -in scf.in > scf.out
#alternatively to re-run a scf calculation copy files from ../phonon/mgb2.save
#mkdir mgb2.save
#cp ../phonon/mgb2.save/charge-density.dat mgb2.save/
#cp ../phonon/mgb2.save/data-file-schema.xml mgb2.save/

ibrun $PATHQE/bin/pw.x -nk 56 -in nscf.in > nscf.out
ibrun $PATHQE/bin/epw.x -nk 56 -in epw2.in > epw2.out

```

```

--                                                                    epw2.in
fbw          = .true.

```

The anisotropic FBW Migdal-Eliashberg equations are solved self-consistently on the imaginary frequency axis by setting the keywords `fbw = .true.`, `eliashberg = .true.`, `laniso = .true.`, and `limag = .true.` in the EPW input file.

The anisotropic FBR Migdal-Eliashberg equations take the following form:

$$\begin{aligned}
 Z_{n\mathbf{k}}(i\omega_j) &= 1 + \frac{T}{\omega_j N_F} \sum_{mj'} \int \frac{d\mathbf{q}}{\Omega_{\text{BZ}}} \frac{\omega_{j'} Z_{m\mathbf{k}+\mathbf{q}}(i\omega_{j'})}{\theta_{m\mathbf{k}+\mathbf{q}}(i\omega_{j'})} \lambda_{n\mathbf{k},m\mathbf{k}+\mathbf{q}}(\omega_j - \omega_{j'}) \\
 \chi_{n\mathbf{k}}(i\omega_j) &= \frac{-T}{N_F} \sum_{mj'} \int \frac{d\mathbf{q}}{\Omega_{\text{BZ}}} \frac{\varepsilon_{m\mathbf{k}'} - \mu + \chi_{m\mathbf{k}+\mathbf{q}}(i\omega_{j'})}{\theta_{m\mathbf{k}+\mathbf{q}}(i\omega_{j'})} \lambda_{n\mathbf{k},m\mathbf{k}+\mathbf{q}}(\omega_j - \omega_{j'}) \\
 \phi_{n\mathbf{k}}(i\omega_j) &= \frac{T}{N_F} \sum_{mj'} \int \frac{d\mathbf{q}}{\Omega_{\text{BZ}}} \frac{\phi_{m\mathbf{k}'}(i\omega_{j'})}{\theta_{m\mathbf{k}+\mathbf{q}}(i\omega_{j'})} [\lambda_{n\mathbf{k},m\mathbf{k}+\mathbf{q}}(\omega_j - \omega_{j'}) - \mu_c^*] \quad (11)
 \end{aligned}$$

where $\lambda_{n\mathbf{k},m\mathbf{k}+\mathbf{q}}(\omega_j - \omega_{j'})$ is the anisotropic electron-phonon coupling strength. The superconducting gap is defined in terms of the renormalization function and the order parameter as: $\Delta_{n\mathbf{k}}(i\omega_j) = \phi_{n\mathbf{k}}(i\omega_j)/Z_{n\mathbf{k}}(i\omega_j)$. The semiempirical Coulomb parameter μ_c^* is provided as an input variable `muc`.

This set of equations is supplemented with an equation for the electron number n_e which determines the chemical potential μ if `muchem = .true.` is set in the EPW calculation.

$$n_e = 1 - \frac{2T}{M} \sum_{n\mathbf{k}} \sum_j \frac{\varepsilon_{n\mathbf{k}} - \mu + \chi_{n\mathbf{k}}(i\omega_j)}{\theta_{n\mathbf{k}}(i\omega_j)} \quad (12)$$

Here, n_e is the average number of electrons per band and M is the total number of momentum-band states.


```

=====
Solve full-bandwidth anisotropic Eliashberg equations
=====
.....
temp( 1) =      10.00000 K

Solve full-bandwidth anisotropic Eliashberg equations on imaginary-axis

Total number of frequency points nsiw( 1) =      92
Cutoff frequency wscut =      0.5008

Size of allocated memory per pool: ~=      0.0680 Gb
iter      ethr      znormi      deltai [meV]      shifti [meV]      mu [eV]
1  2.863890E+00  1.503340E+00  6.227401E+00  7.559645E-02  7.469369E+00
2  7.142968E-02  1.471850E+00  6.504459E+00  4.574388E-01  7.469369E+00
.....
9  6.092146E-04  1.458403E+00  6.920549E+00  4.458350E-01  7.469369E+00
Convergence was reached in nsiter =      9

Chemical potential (itemp = 1) =      7.4693693827E+00 eV

Temp (itemp = 1) =      10.000 K Free energy =      -0.010531 meV

Min. / Max. values of superconducting gap =      0.000000      13.394847 meV

```

► 11th step: To plot the various physical quantities saved in the output files, follow the steps 4 to 7 above. You can also compare your results with those from the previous FSR calculation.

How to plot the superconducting gap on the Fermi surface with VESTA:

1. Plot Fermi surface (FS)

mgb2.fs_YY.cube (YY = band index within the [fsthick](#)) files were generated by setting `fermi_plot = .true.` in `epw1.in`. Each file contains the energy eigenvalues relative to the Fermi level, and can be visualized with [VESTA](#).

To visualize, open `mgb2.fs_1.cube` with VESTA and then import `mgb2.fs_2.cube`, ... files one-by-one as follows:

Edit → Edit Data → Volumetric Data → Import (under Isosurface) → Choose: Multiply to current data

Uncheck: Style → Structural Models → Show models

Set: Properties → Isosurfaces → Isosurface level: 0

Set: Properties → Sections → Opacity of drawn sections(%): 0

2. Color the FS based on the superconducting gap values at a specific temperature (e.g., 10.0 K)

Import: `mgb2.imag_aniso_gap0.010.00_1.cube`, `mgb2.imag_aniso_gap0.010.00_2.cube`, ... files one-by-one on FS as the following;

Edit → Edit Data → Volumetric Data → Import (under Surface coloring) → Choose: Add to current data

You should get the following plot at 10 K

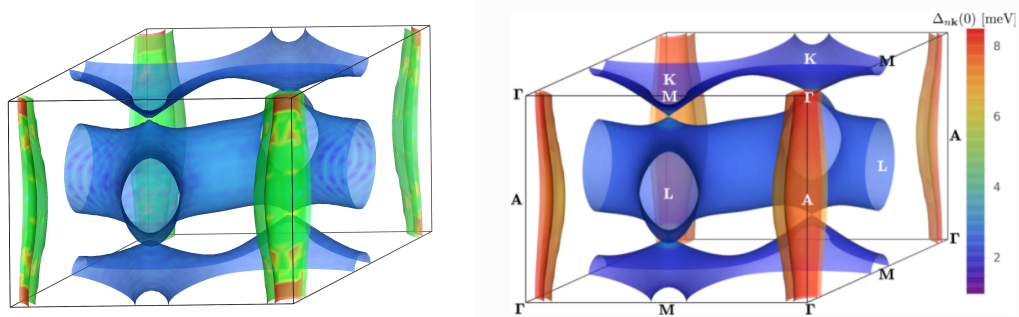


Fig. 11 Calculated superconducting gap of MgB_2 on the Fermi surface at 10 K. At convergence you should get the right hand-side figure adapted from [Comp. Phys. Comm. 209, 116 \(2016\)](#).

How to plot the superconducting gap on the Fermi surface with FermiSurfer:

To visualize, open `mgb2.imag_aniso_gap0_XX.frmsf` with FermiSurfer.

`mgb2.imag_aniso_gap0_XX.frmsf` ($XX = \text{temperature}$) files were generated by setting `laniso = .true.` in `epw1.in`. Each file contains the energy eigenvalues relative to the Fermi level, and can be visualized with [FermiSurfer](#).

Notes on input variables:

- `ephwrite = .true.` does not work with random \mathbf{k} or \mathbf{q} grids and requires `nkf1`, `nkf2`, `nkf3` to be multiple of `nqf1`, `nqf2`, `nqf3`.
- `mp_mesh_k = .true.` specifies that only the irreducible points for the dense \mathbf{k} grid are used. This significantly decreases the computational cost when solving the anisotropic Migdal-Eliashberg equations.
- If the anisotropic Migdal-Eliashberg equations are solved in a separate run from the one in which the `ephmatXX`, `freq`, `egnv`, and `ikmap` files saved in `prefix.ephmat` directory were generated, the code requires to use **the same number of CPUs as the number of ephmatXX files**. If you forget this the code will stop with a message asking to use `npool` equal to the number of `ephmatXX` files.
- `lpade = .true.` requires `limag = .true.`
- `lacon = .true.` requires both `limag = .true.` and `lpade = .true.`
- `wscut` gives the upper limit (in eV) of the summation over the Matsubara frequencies on the imaginary axis in the Migdal-Eliashberg equations (`limag = .true.`). Note that the input variable `wscut` is ignored if the number of frequency points is given using the input variable `nswi`. In this case, the number of frequency points in the summation is the same irrespective of the temperature.
- `temps = t1 t2 t3 ...` define the list of temperatures at which the Migdal-Eliashberg equations are evaluated. Note that an evenly spaced temperature grids can also be defined using `nstemp`, `temps = min.temp max.temp` input variables.
- If temperatures larger than the critical temperature T_c estimated using the Allen-Dynes formula are specified in the input file a warning message is written in the output file. The code may stop when such a temperature is reached if the Migdal-Eliashberg equations do not have a solution at that point.

- `muchem` solve the anisotropic FBW ME eqs. with variable chemical potential.
- `gridsamp = 0` generates a uniform Matsubara frequency grid (default).
- `gridsamp = 1` generates a sparse Matsubara frequency grid.
- `imag_read` works if `limag = .true.` and `laniso = .true.` and it allows the code to read from file the superconducting gap and renormalization function on the imaginary axis at specific temperature `XX` from file `prefix.imag_aniso_XX`. The temperature is specified as `temps = XX` (first temperature) in the EPW input file.
- `imag_read` can be used to: (1) solve the anisotropic Migdal-Eliashberg equations on the imaginary axis at temperatures greater than `XX` using as a starting point the superconducting gap estimated at temperature `XX`. (2) obtain the solutions of the anisotropic Migdal-Eliashberg equations on the real axis with `lpade = .true.` or `lacon = .true.` starting from the imaginary axis solutions at temperature `XX`; (3) write to file the anisotropic superconducting gap on the Fermi surface in cube format at temperature `XX` for `verbosity = 2`. The generated output files are `prefix.imag_aniso_gap_XX_YY.cube`, where `YY` is the band number within the chosen energy window during the EPW calculation.

Restart options (this requires to use the same number of cores as in the original run):

1. Restart from an interrupted `q`-point while writing `ephmatXX` files.

Required files: `prefix.epmatwp`, `prefix.ukk`, `crystal.fmt`, `epwdata.fmt`, `vmedata.fmt` (or `dmedata.fmt`), `restart.fmt`, and `selecq.fmt` (`selecq.fmt` only needed if `selecqread = .true.` otherwise it will be re-created).

Input setup:

```
ep_coupling = .true.
elph         = .true.

epwwrite    = .false.
epwread     = .true.           ! read *.epmatwp and *.fmt files

wannierize  = .false.         ! read *.ukk file
ephwrite    = .true.
```

2. Restart by reading `ephmatXX` files.

Required files: `prefix.ephmat` directory (which contains `egnv`, `freq`, `ikmap`, `ephmatXX` files), `selecq.fmt`, and `crystal.fmt`

Input setup:

```
ep_coupling = .false.
elph         = .false.

epwwrite    = .false.
epwread     = .true.

wannierize  = .false.
ephwrite    = .false.
```